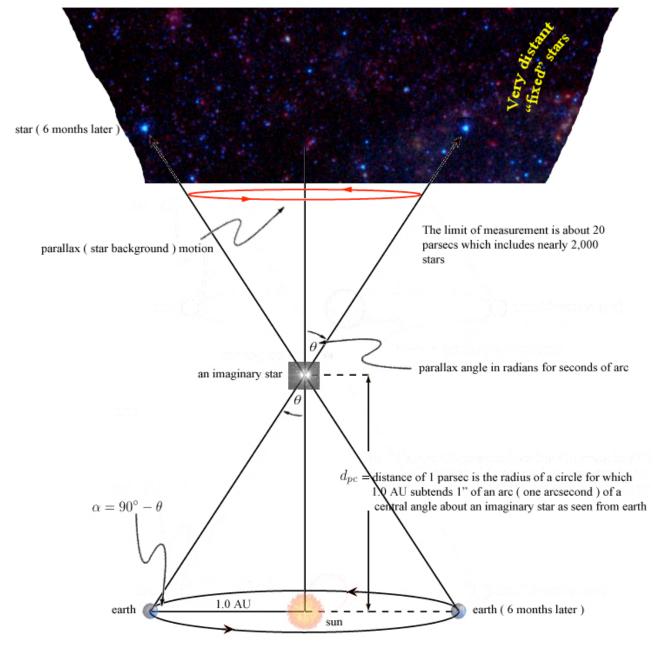
# Structure and Evolution of Stars

Lecture 2







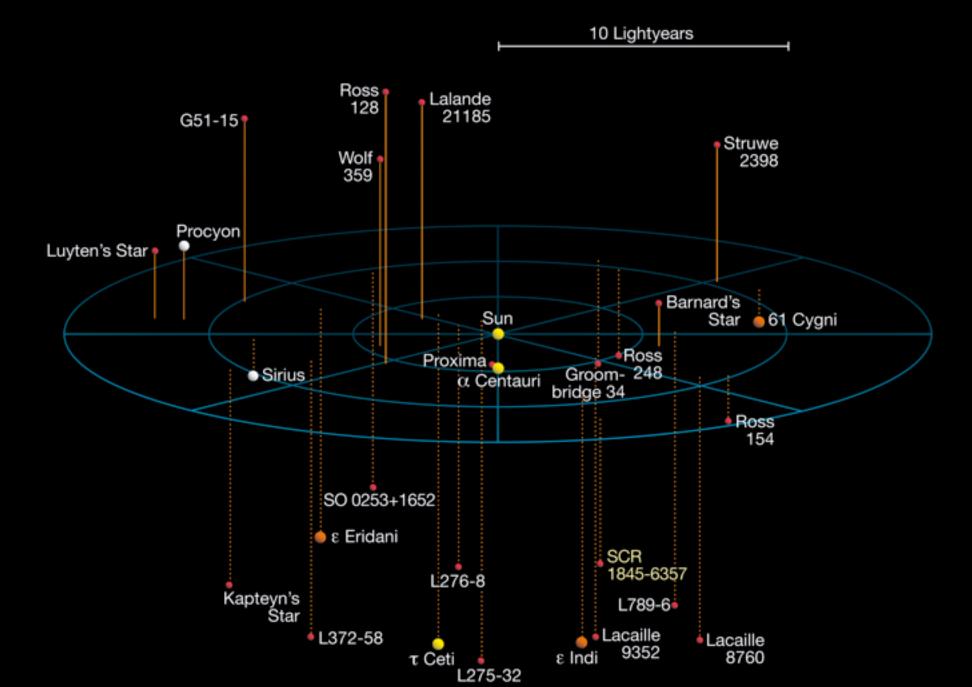


Where:

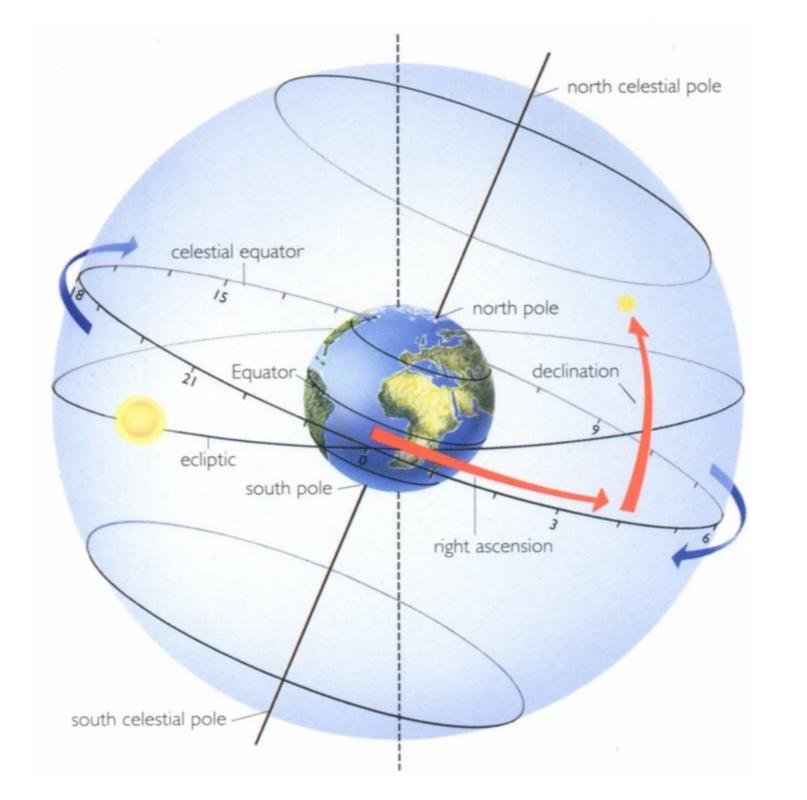
 $\theta$  = angle of parallax in radians for seconds of arc  $d_{P^c}$  = 1/ $\theta$  , distance to an imaginary star in parsecs and is the radius of a circle for which

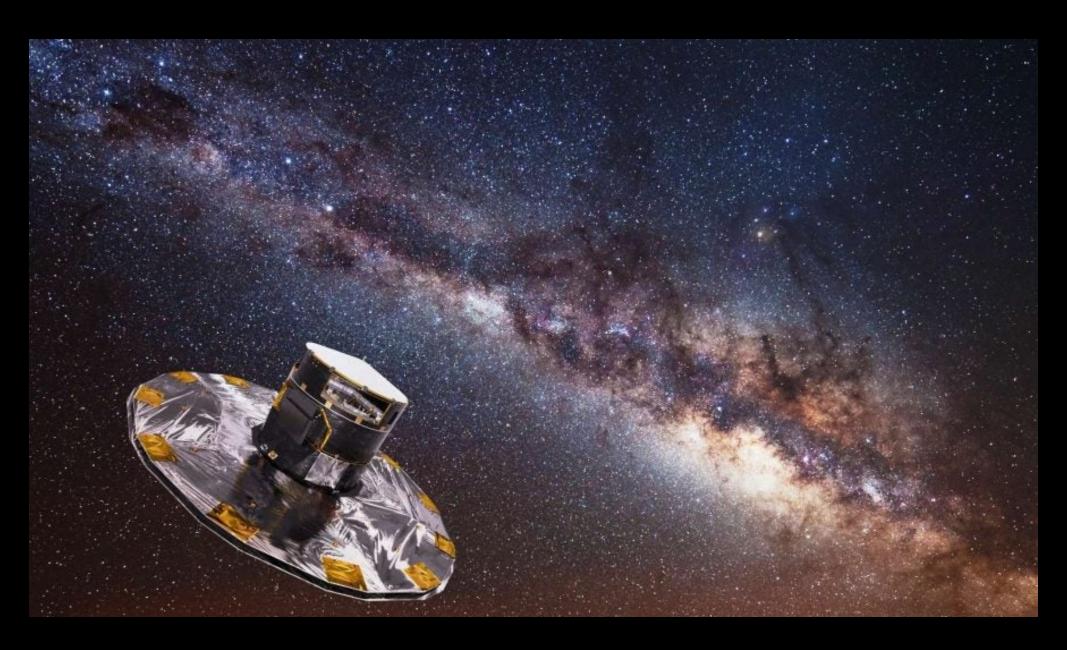
1.0 AU subtends 1.0" (one second) of arc of a central angle about an imaginary star as seen from earth

note: the word parsec stands for "Parallax of one arcsecond"





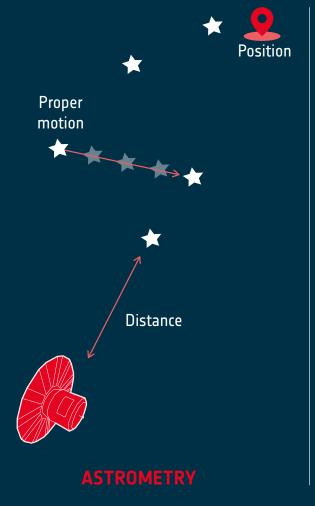


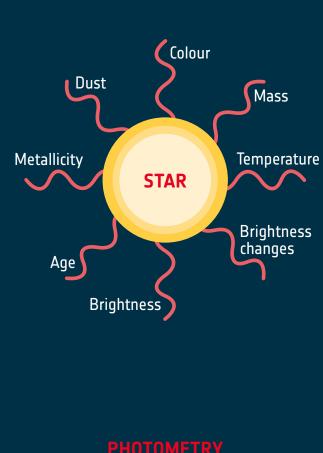


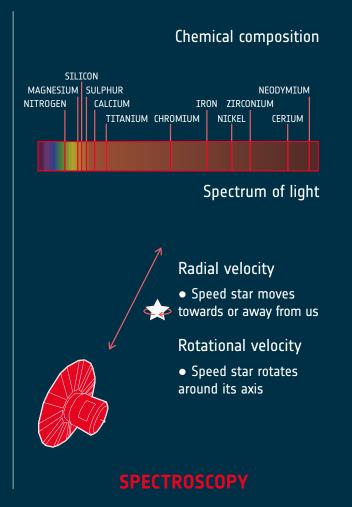
### **GAIA'S OBSERVING TECHNIQUES**



Techniques to study the stars in our cosmic neighbourhood.













#### **MILKY WAY STARS**



Data release 3 includes a total of 1.8 billion Milky Way stars – providing astronomers with an unprecedented view of stellar characteristics and their life cycle,

and the galaxy's structure

and evolution.

**Object** classifications 1.5 billion

What type of star is it? Variable stars 10 million

> Changing brightness over time

Low resolution spectroscopy

470 million astrophysical parameters 220 million spectra

> Temperature | Mass Age | Colour Metallicity

**Binary star systems** 813 thousand

Position | Distance Orbit | Mass

1.8 billion stars

**Radial velocity** 33 million

Speed star moves towards or away from us

> Third velocity dimension

**Astrometry and** photometry 1.5 billion

Already released in the last of the last o Brightness and colour Position | Distance **Proper motions** 

**High resolution** spectroscopy

5.6 million astrophysical parameters

2.5 million chemical compositions

1 million spectra

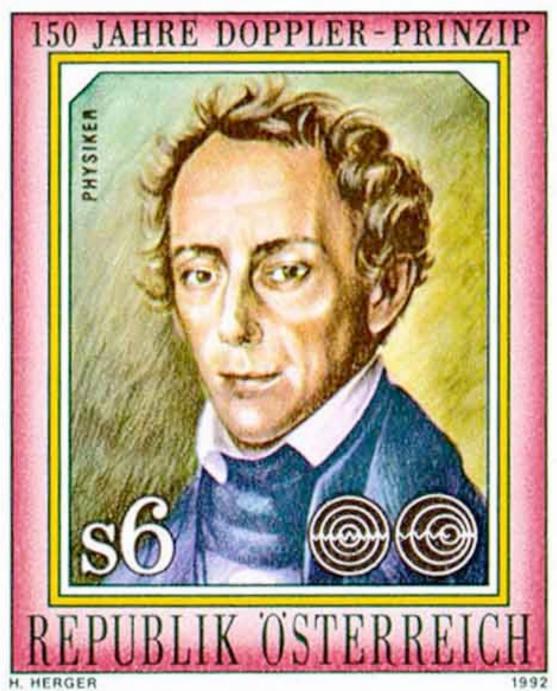
Chemical composition Temperature | Mass | Age



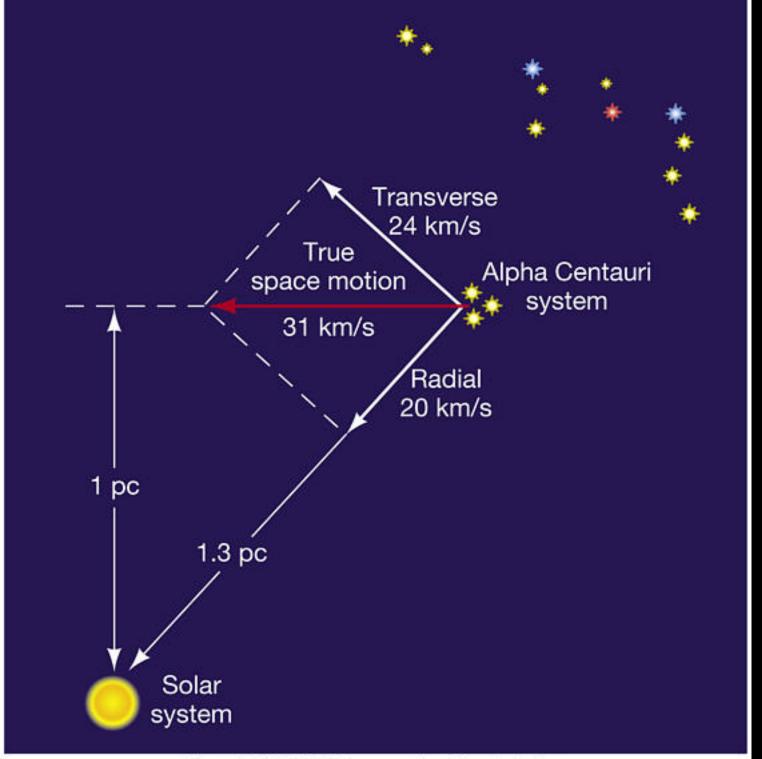




- 2007



H. HERGER



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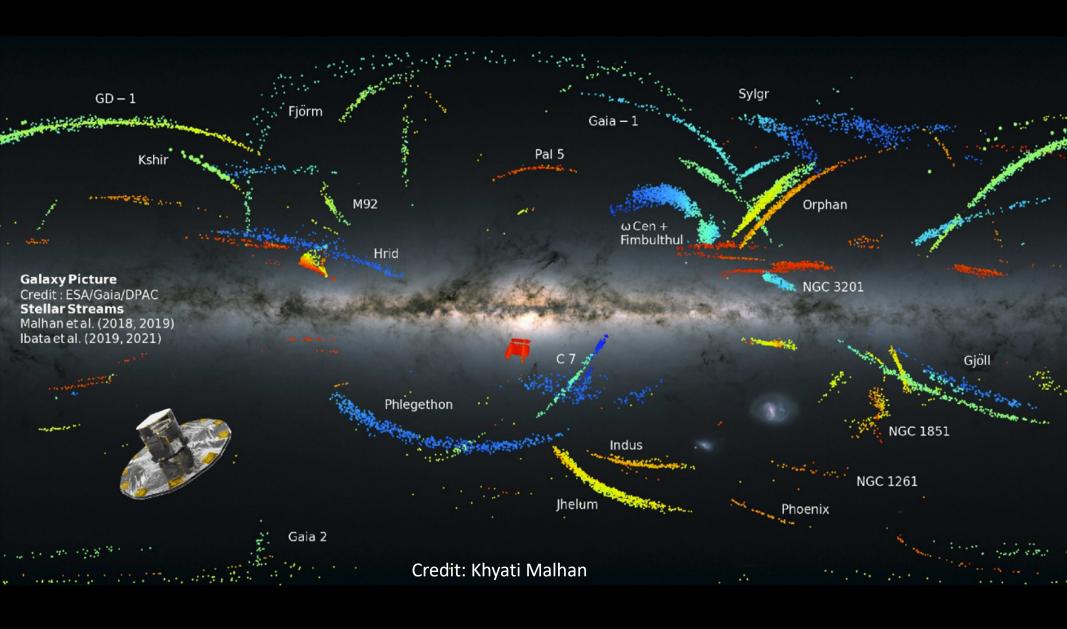
1 million spectra

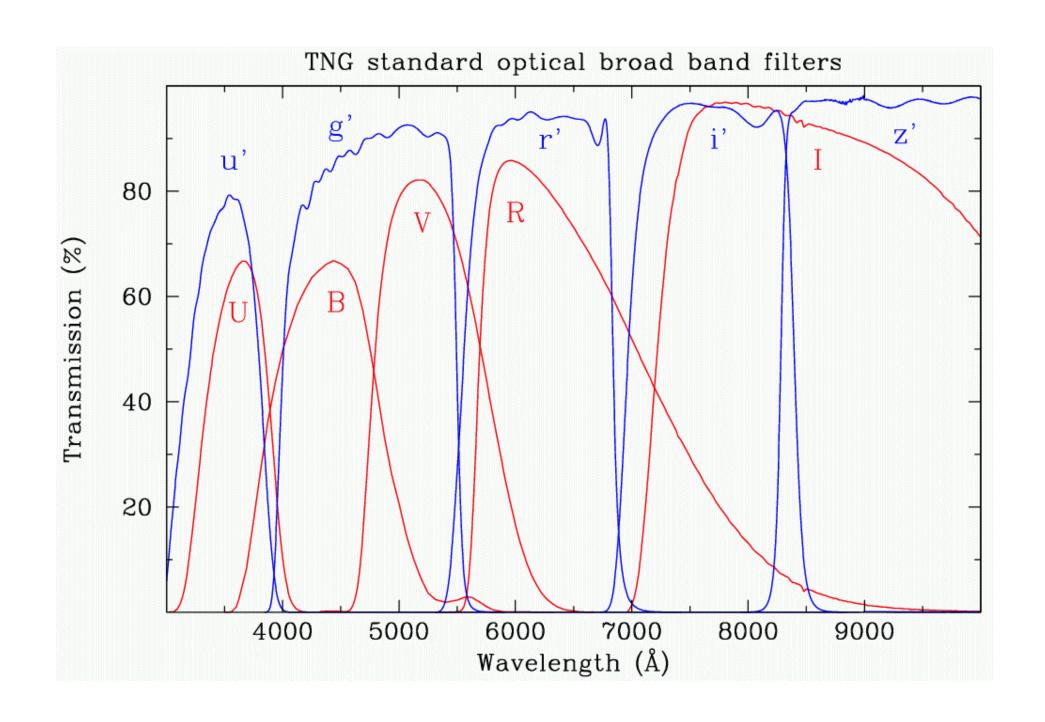
Chemical composition Temperature | Mass | Age



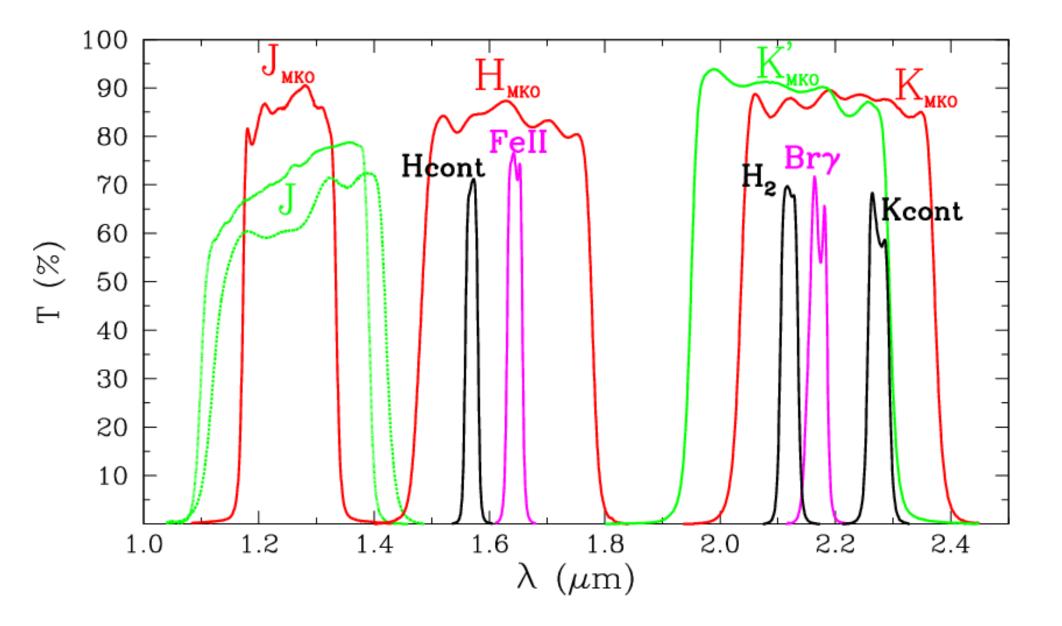


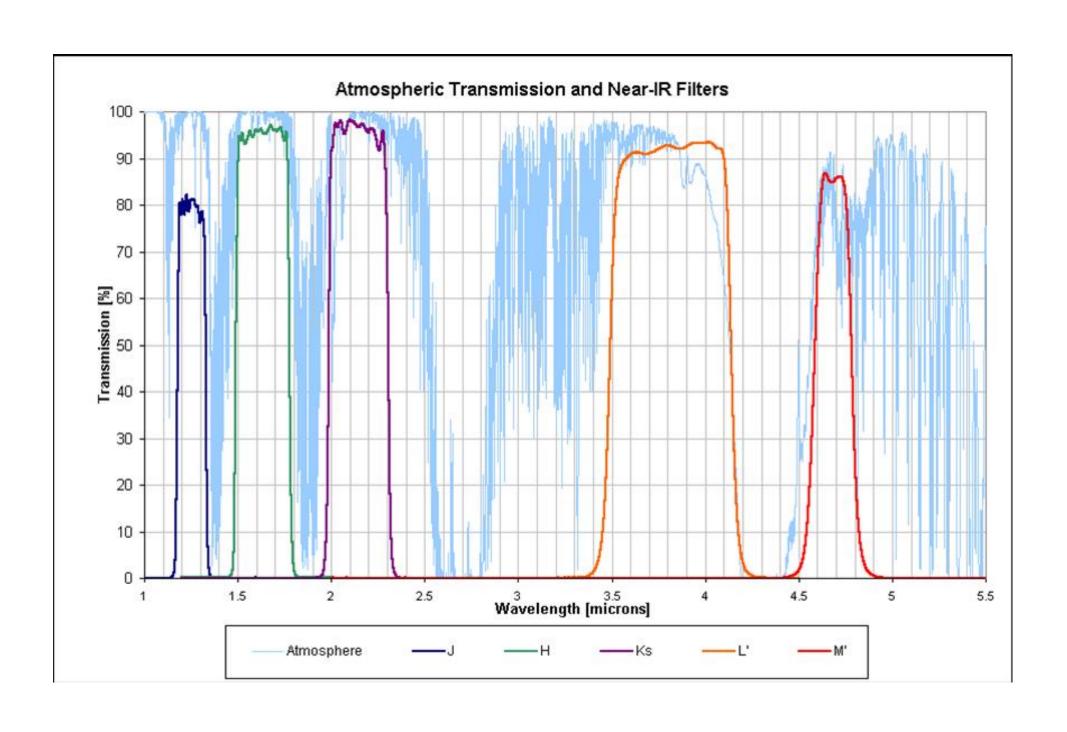




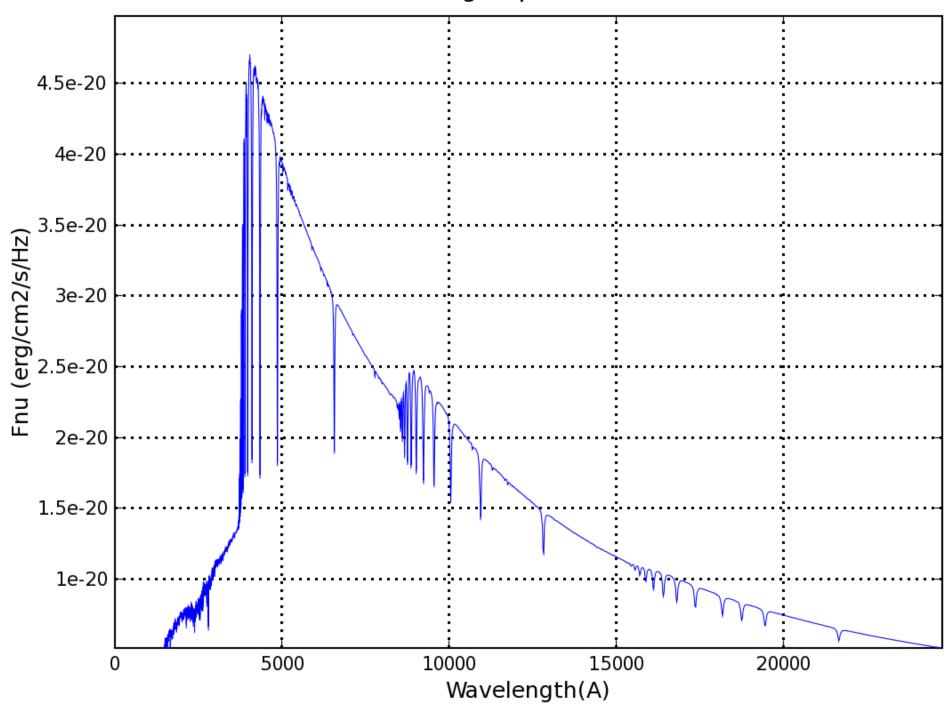


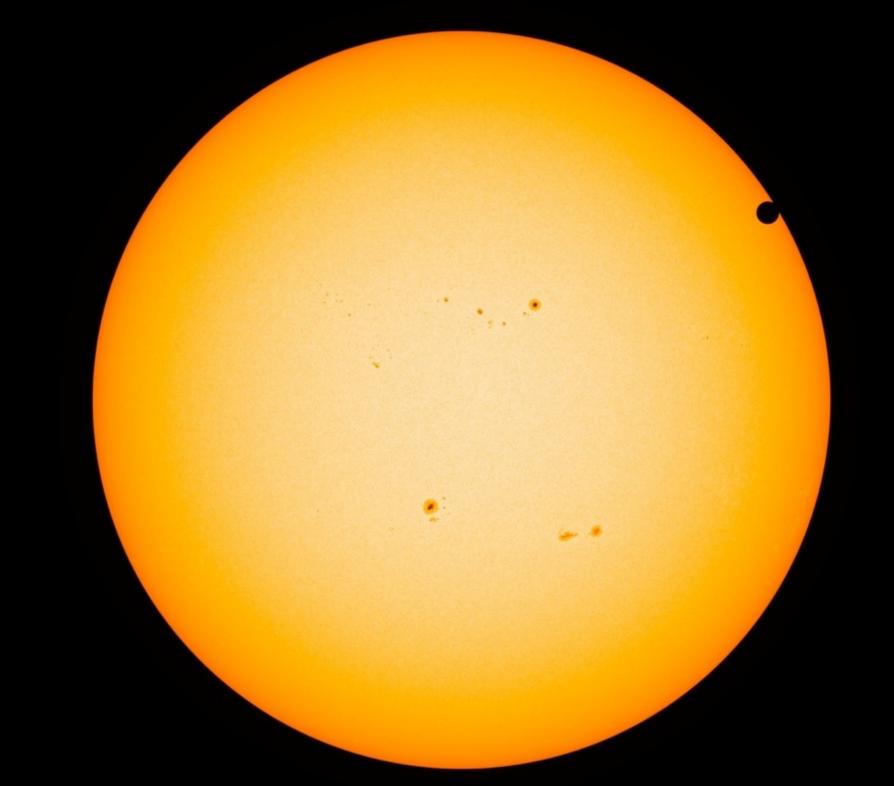


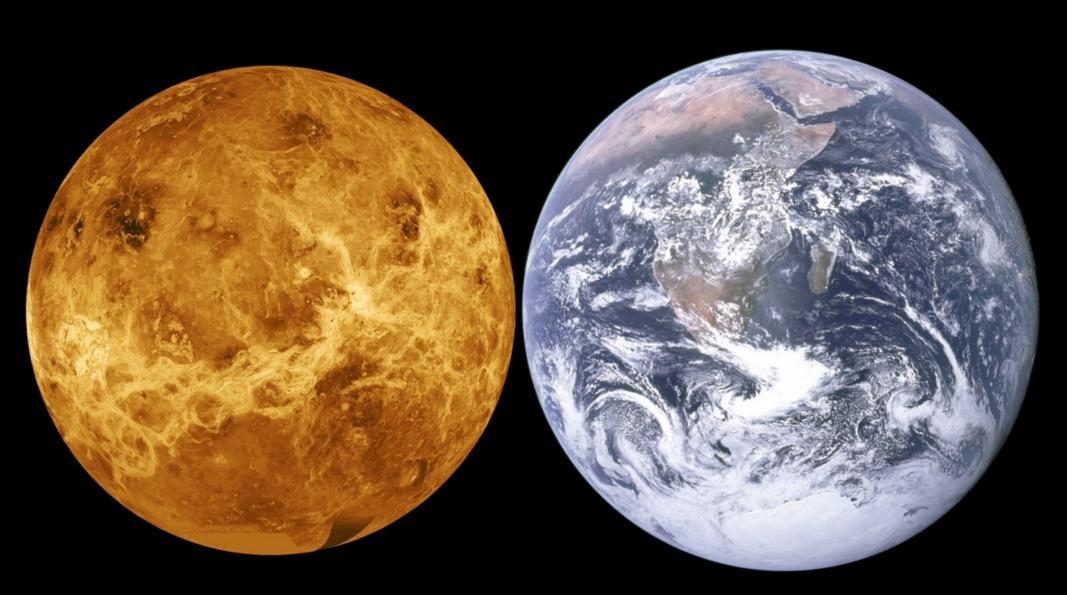




## Vega Spectrum







# Sun's Spectrum vs. Thermal Radiator

of a single temperature T = 5777 K

