



# Structure and Evolution of Stars

## Lecture 2







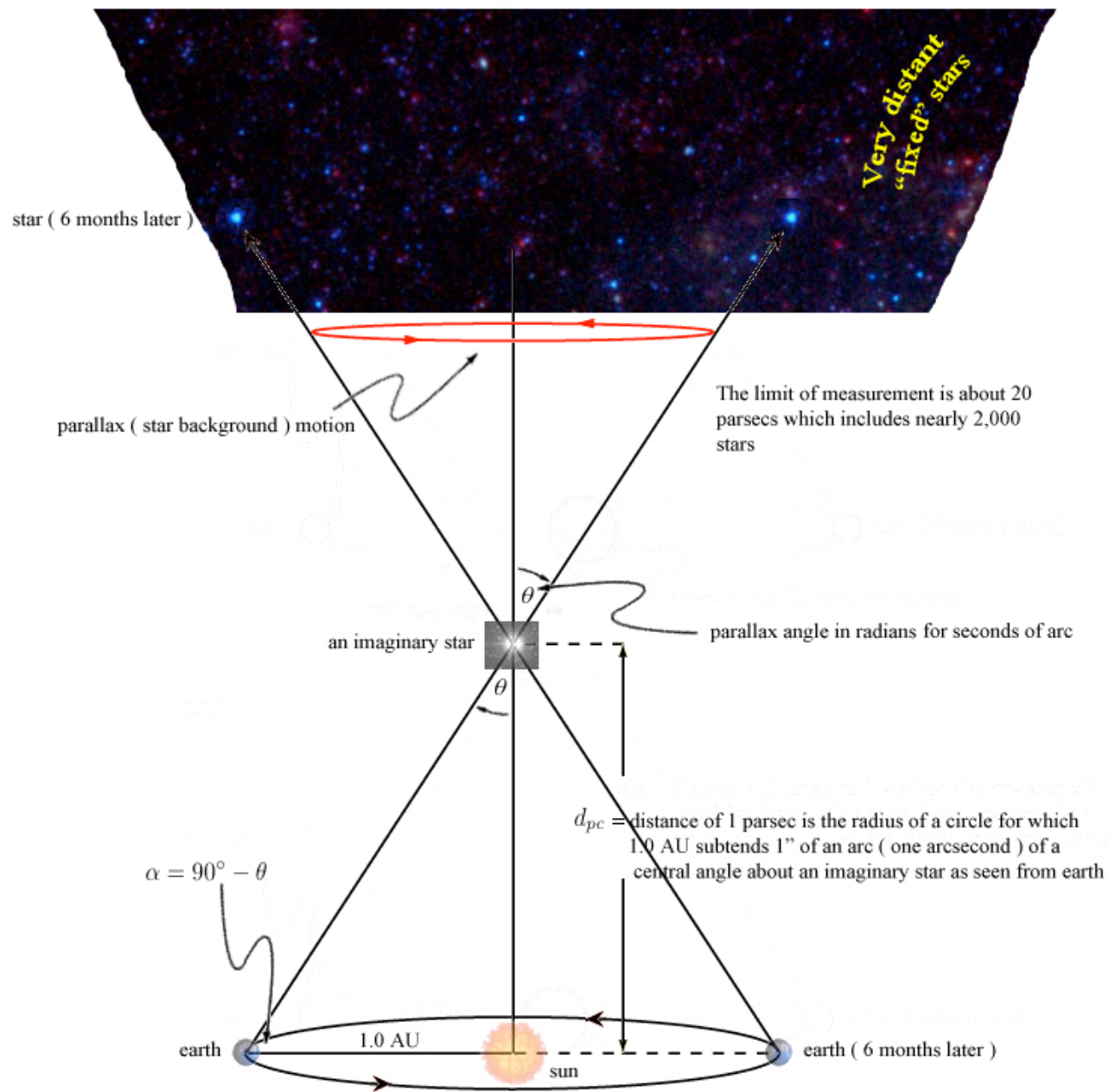
NASA, ESA, CSA, STScI  
Webb ERO Production Team











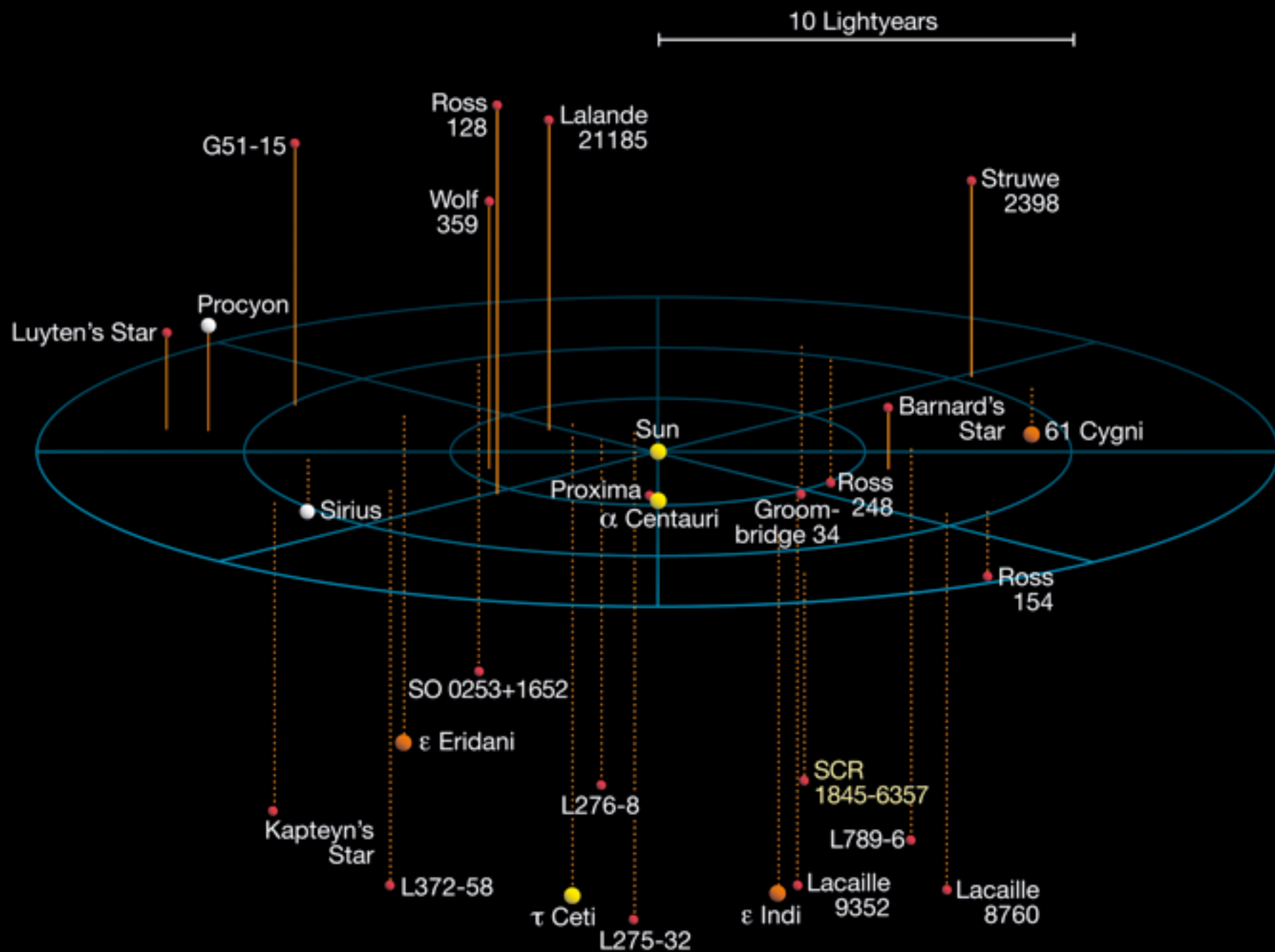
Where:

$\theta$  = angle of parallax in radians for seconds of arc

$d_{pc} = 1/\theta$ , distance to an imaginary star in parsecs and is the radius of a circle for which 1.0 AU subtends 1.0" ( one second ) of arc of a central angle about an imaginary star as seen from earth

note: the word parsec stands for "Parallax of one arcsecond"

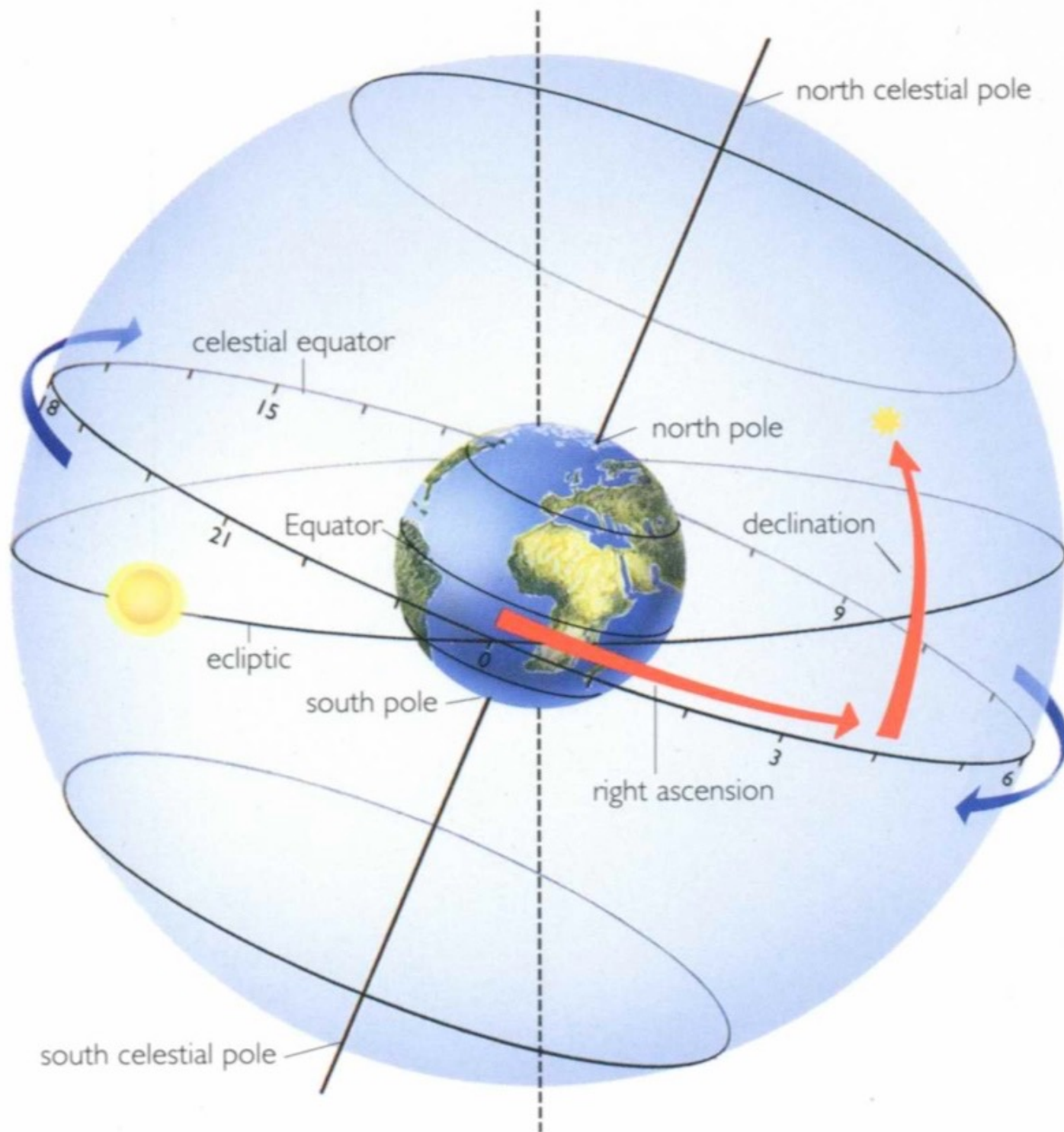




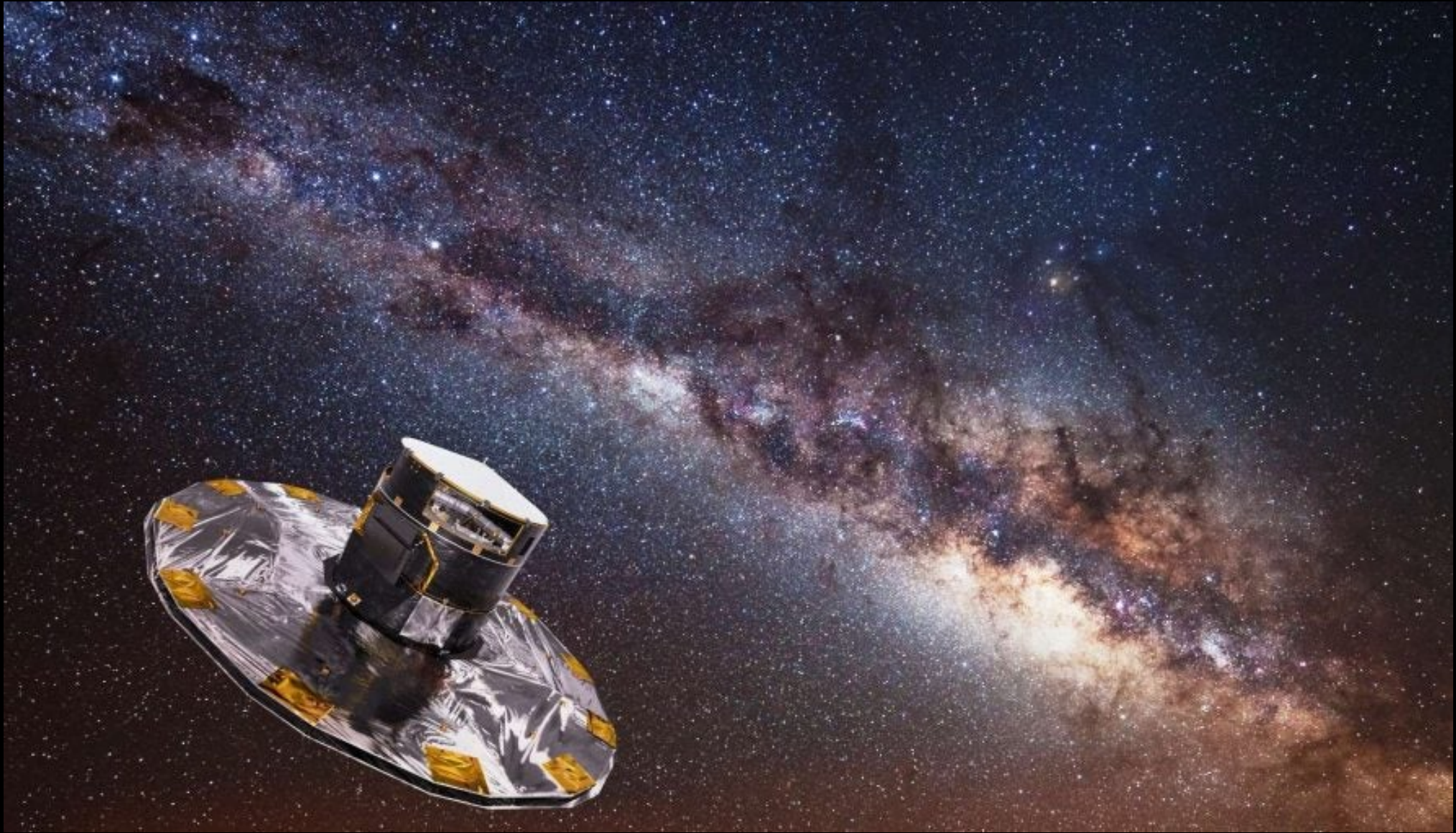










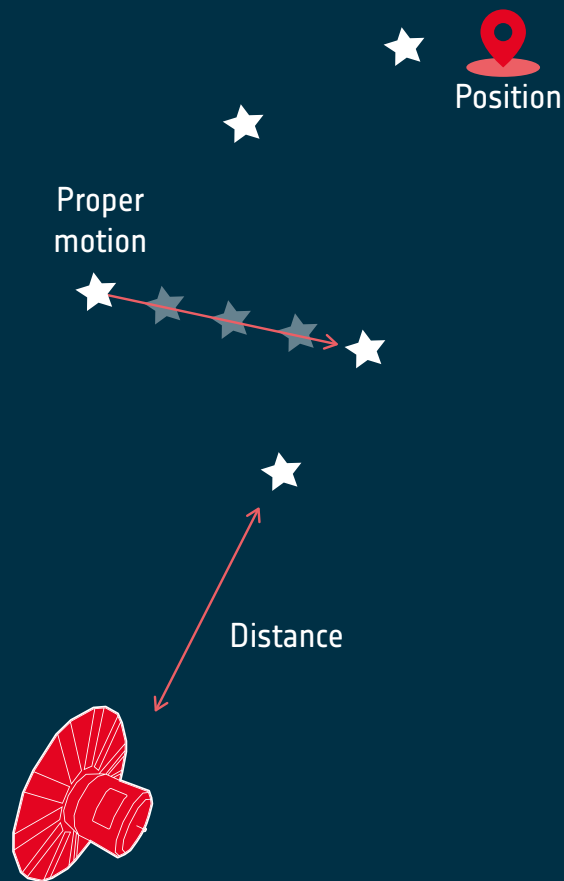




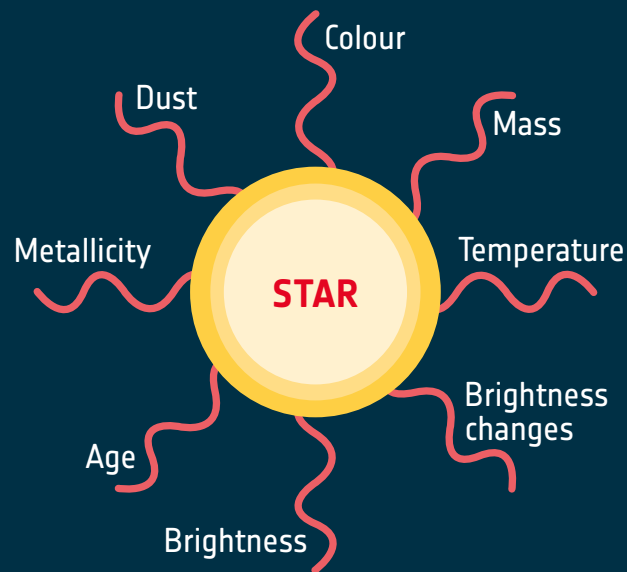
# GAIA'S OBSERVING TECHNIQUES



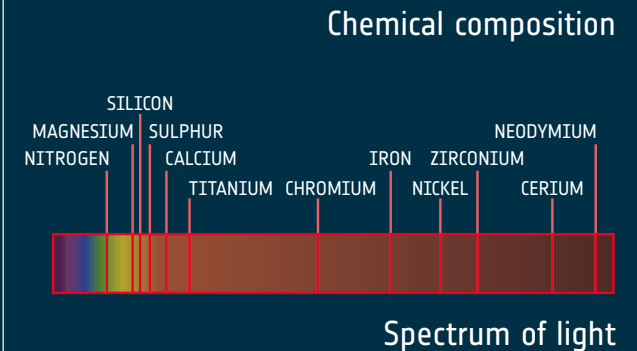
Techniques to study the stars in our cosmic neighbourhood.



**ASTROMETRY**



**PHOTOMETRY**



**SPECTROSCOPY**

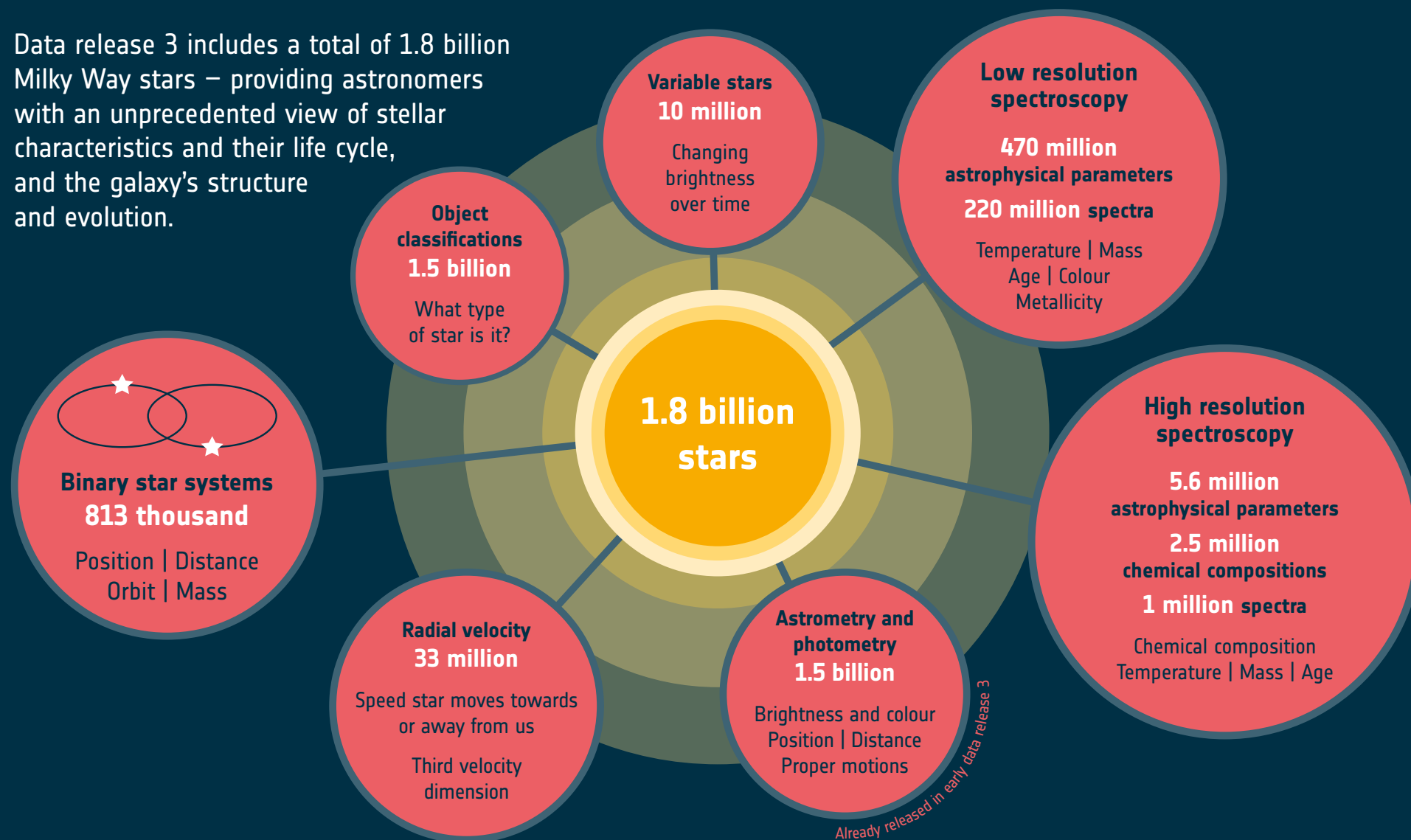




# MILKY WAY STARS



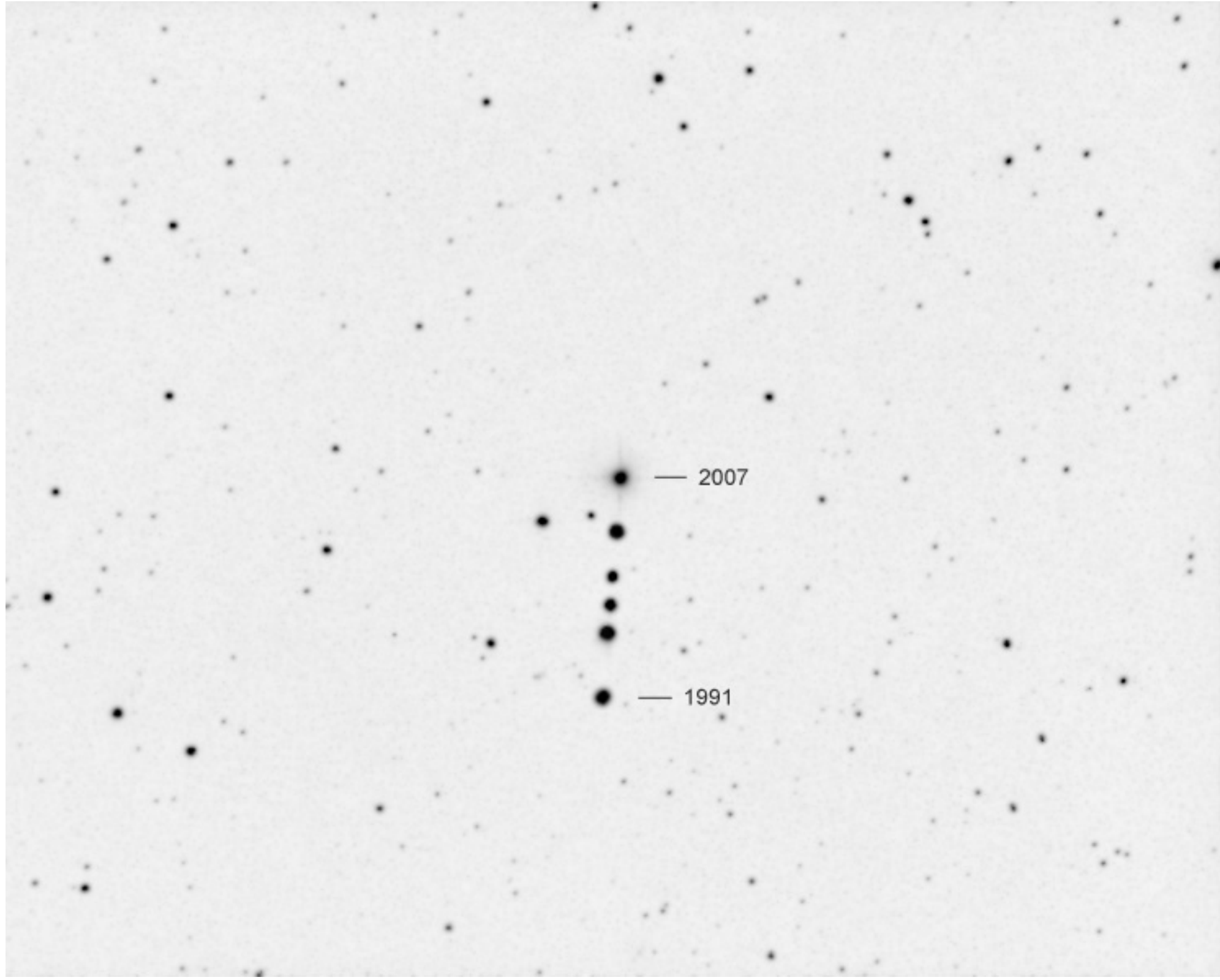
Data release 3 includes a total of 1.8 billion Milky Way stars – providing astronomers with an unprecedented view of stellar characteristics and their life cycle, and the galaxy's structure and evolution.

















150 JAHRE DOPPLER-PRINZIP

PHYSIKER

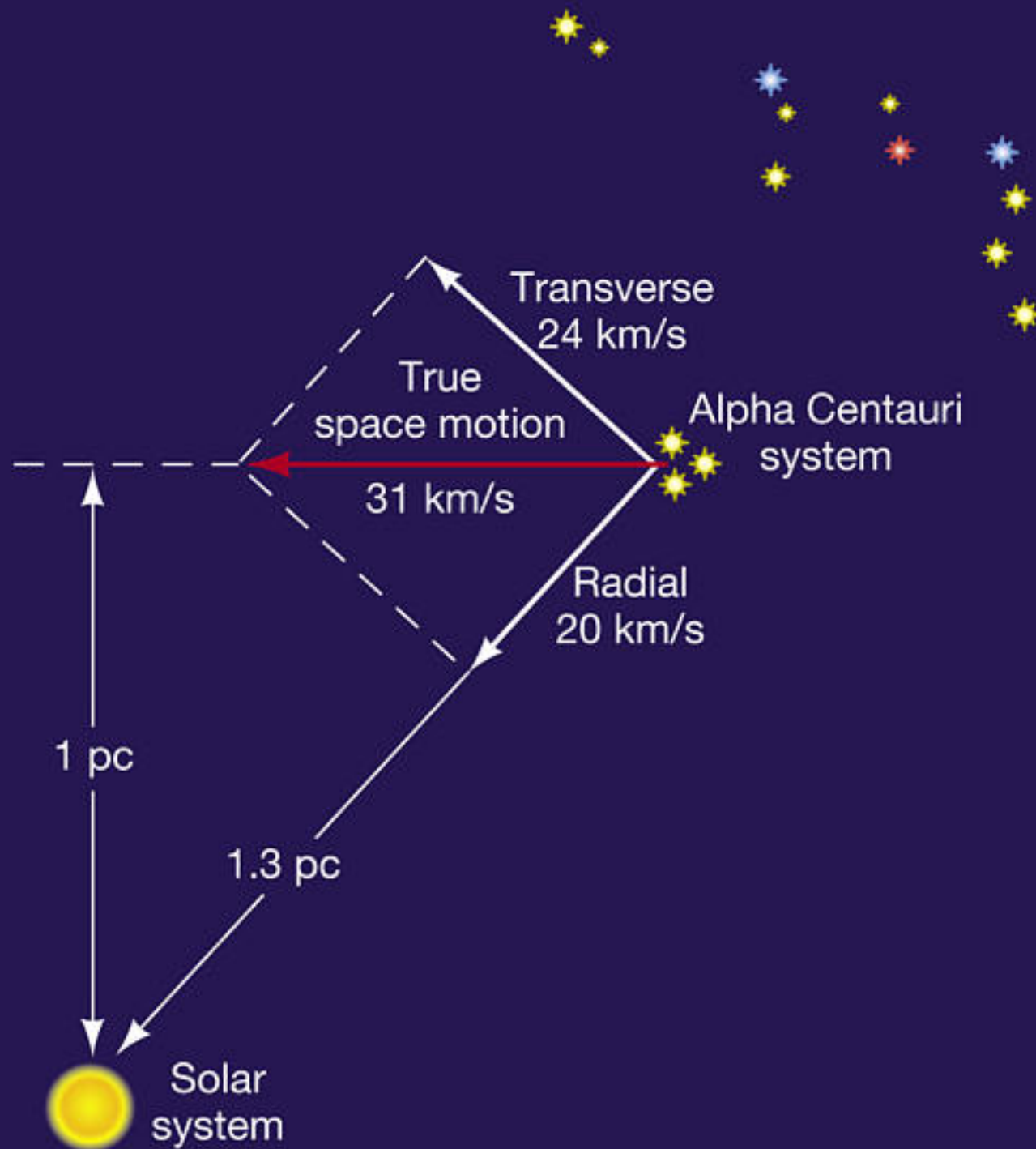


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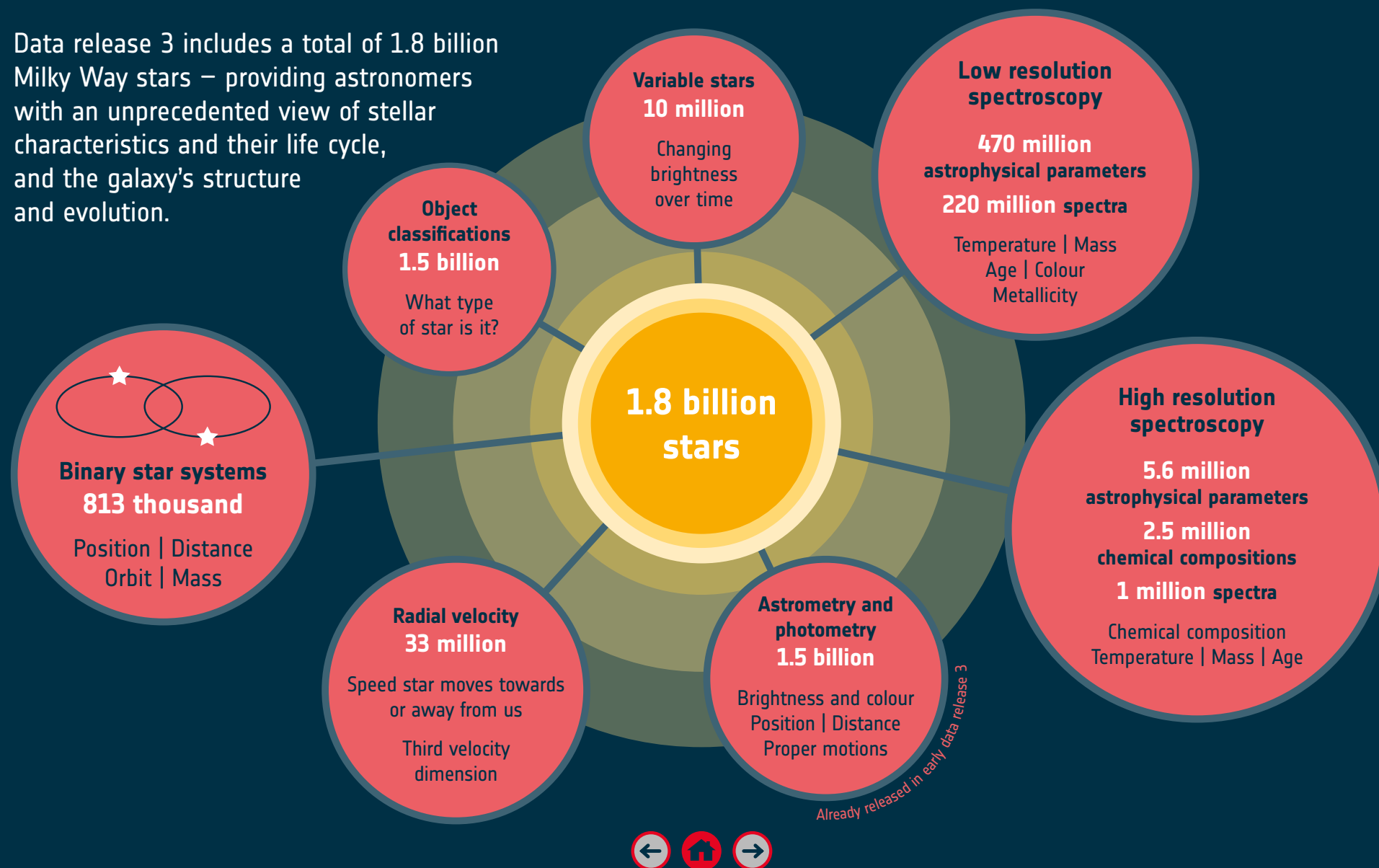




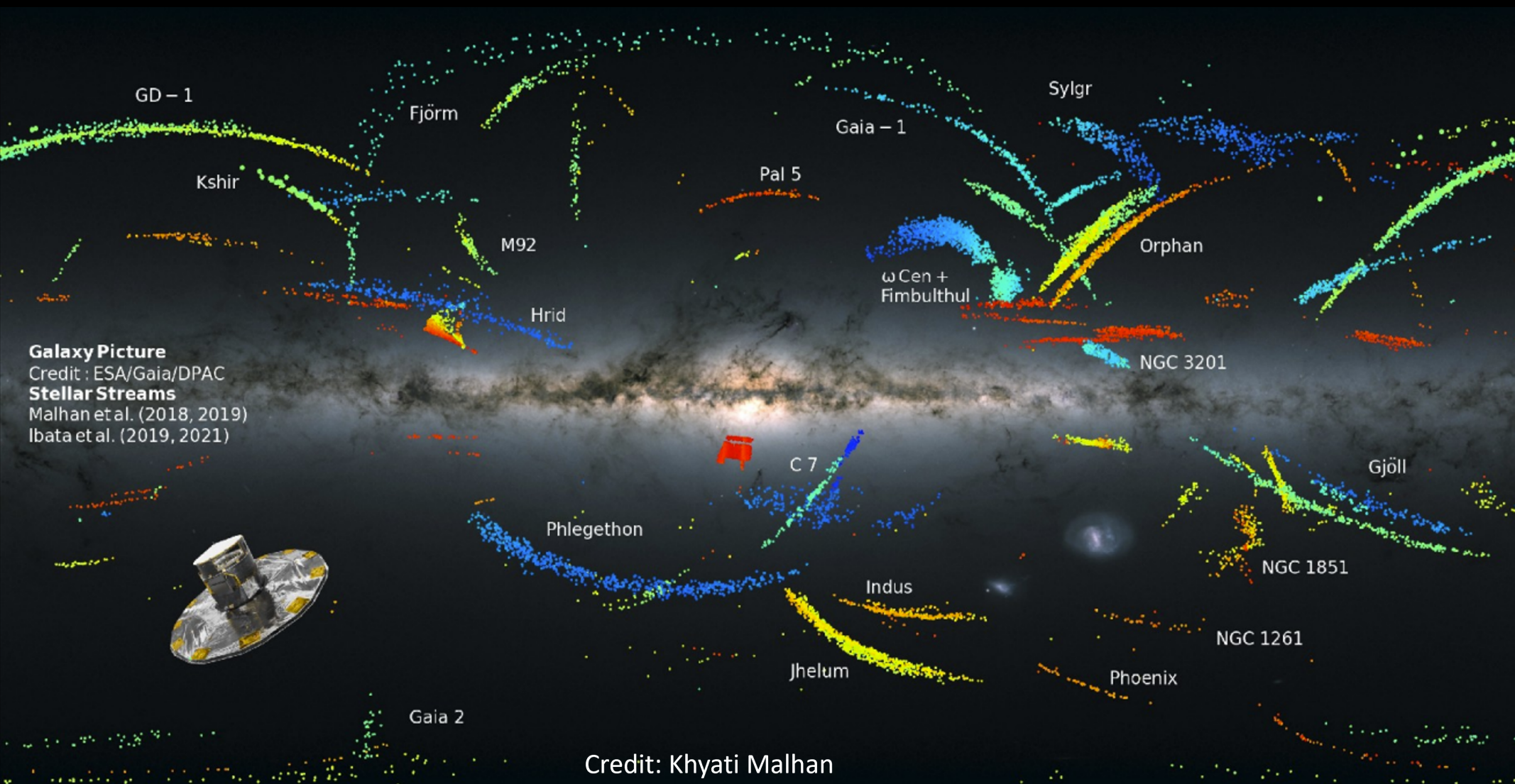
# MILKY WAY STARS



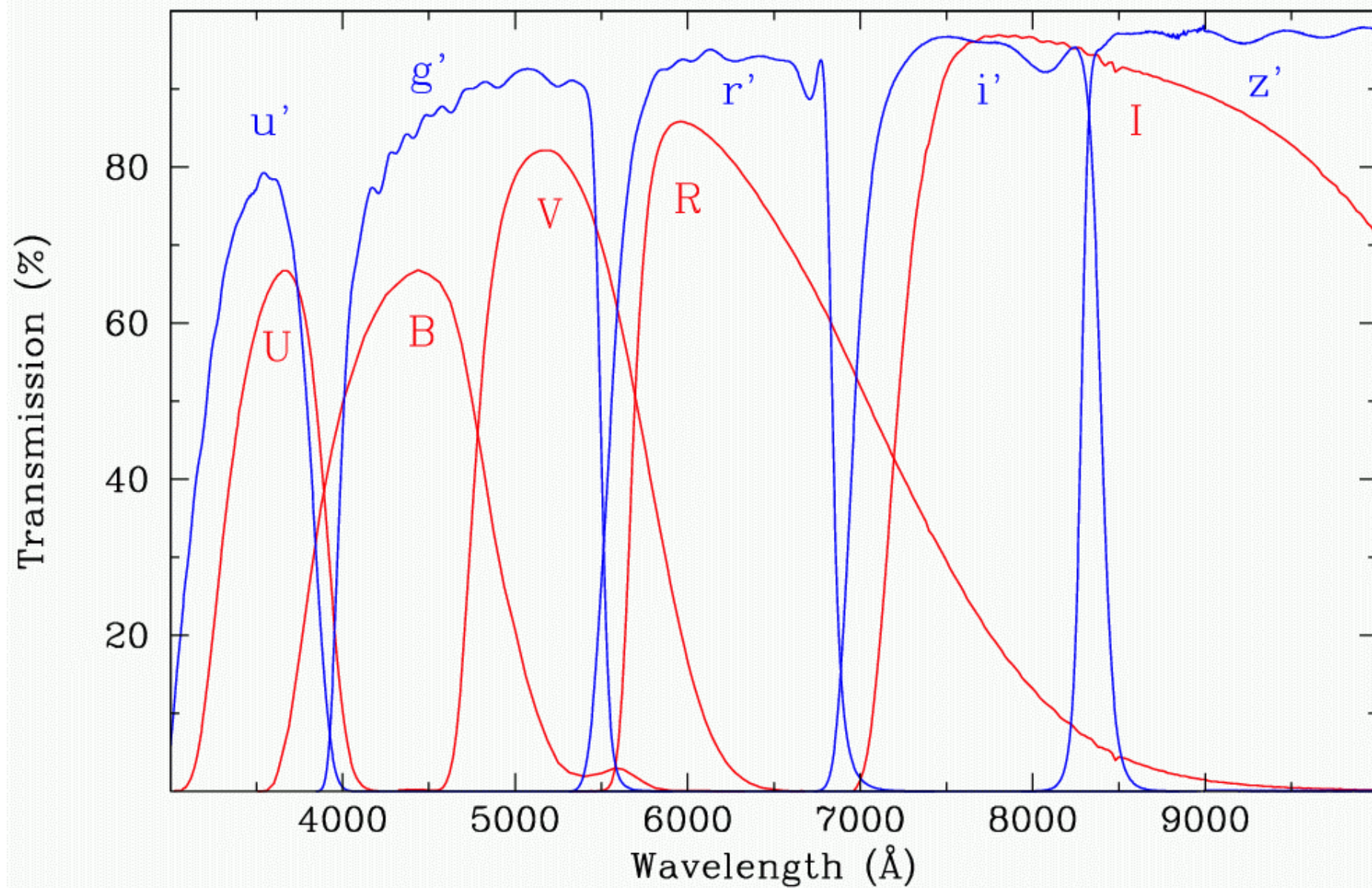
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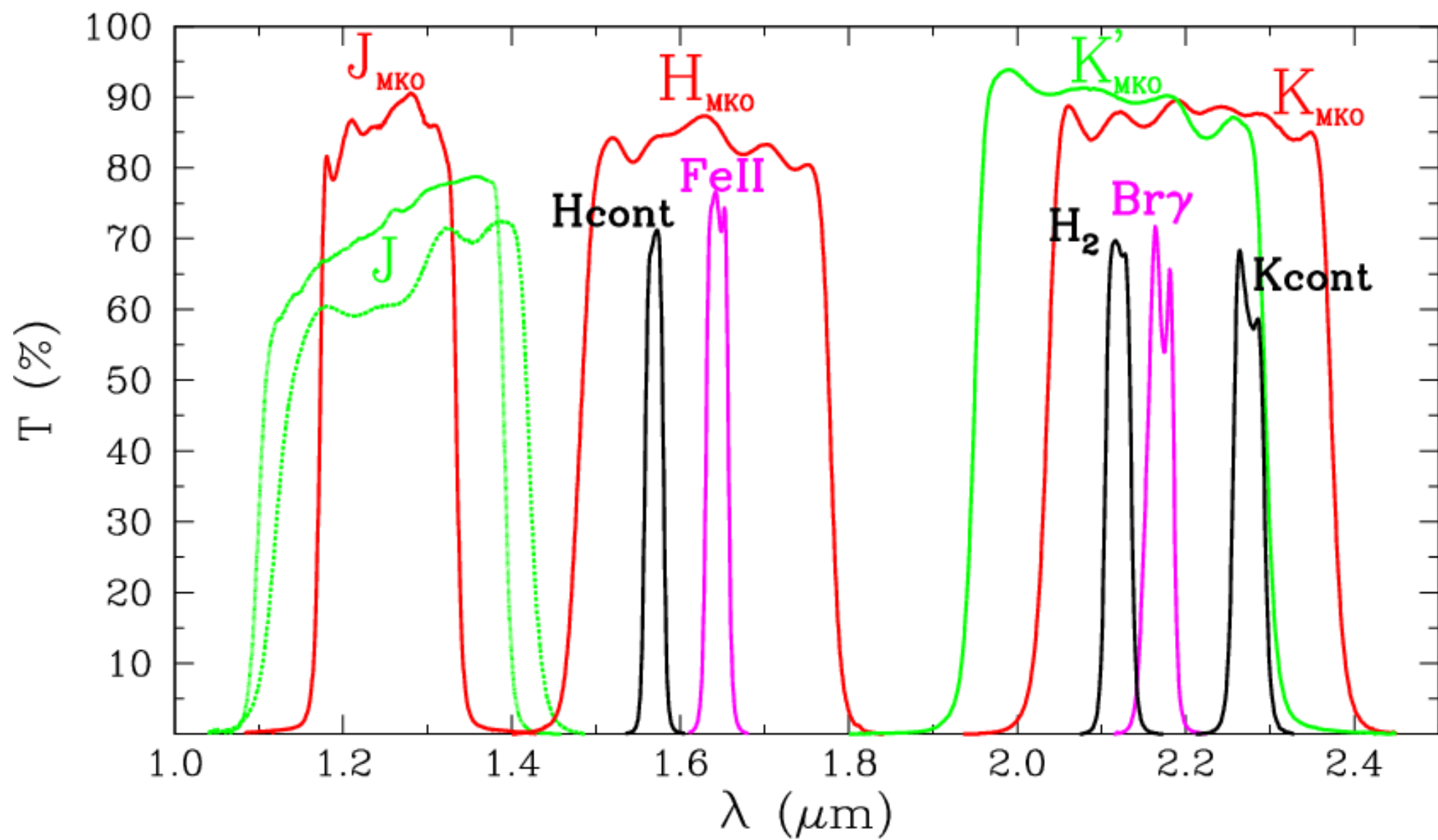


TNG standard optical broad band filters



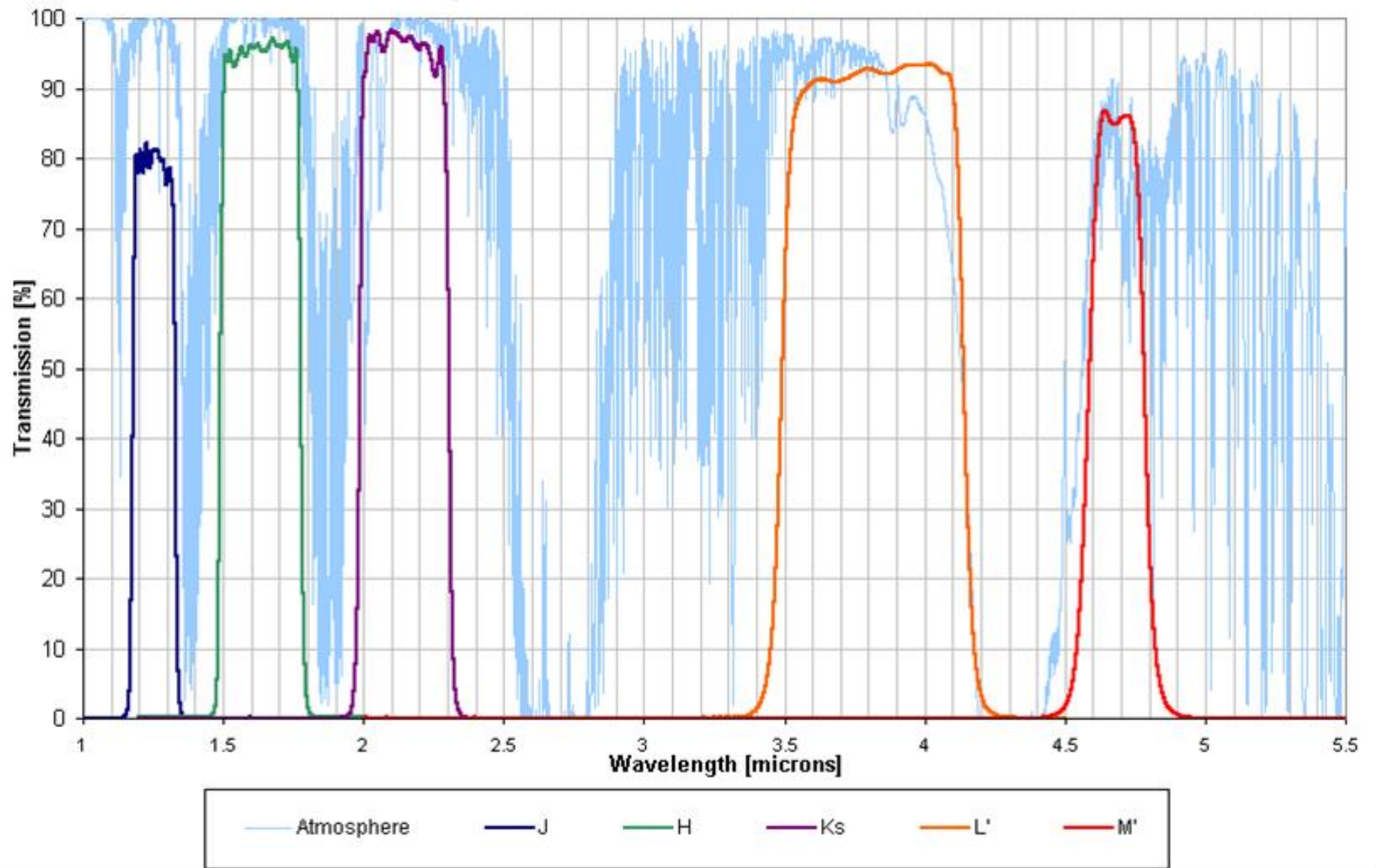




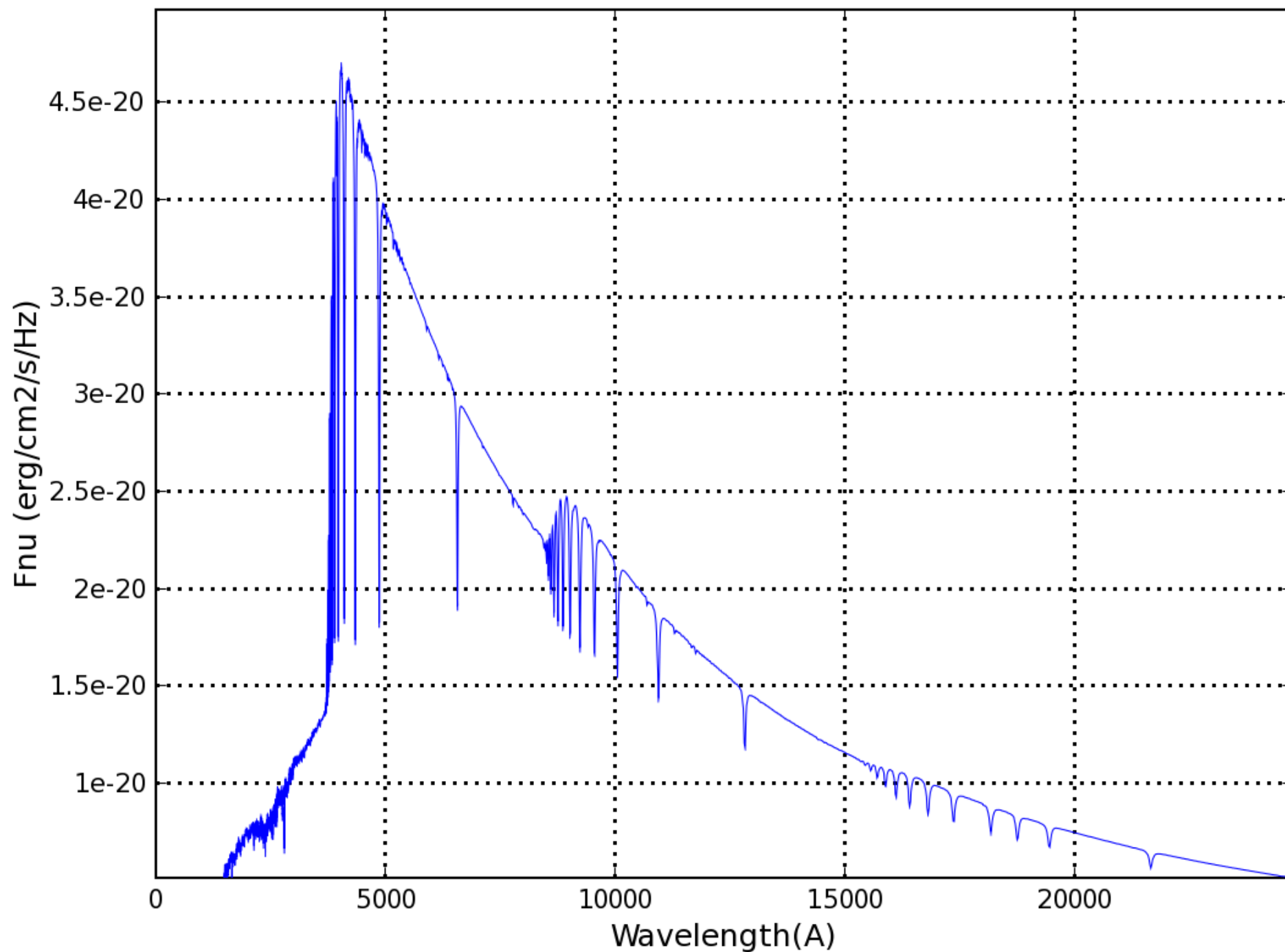




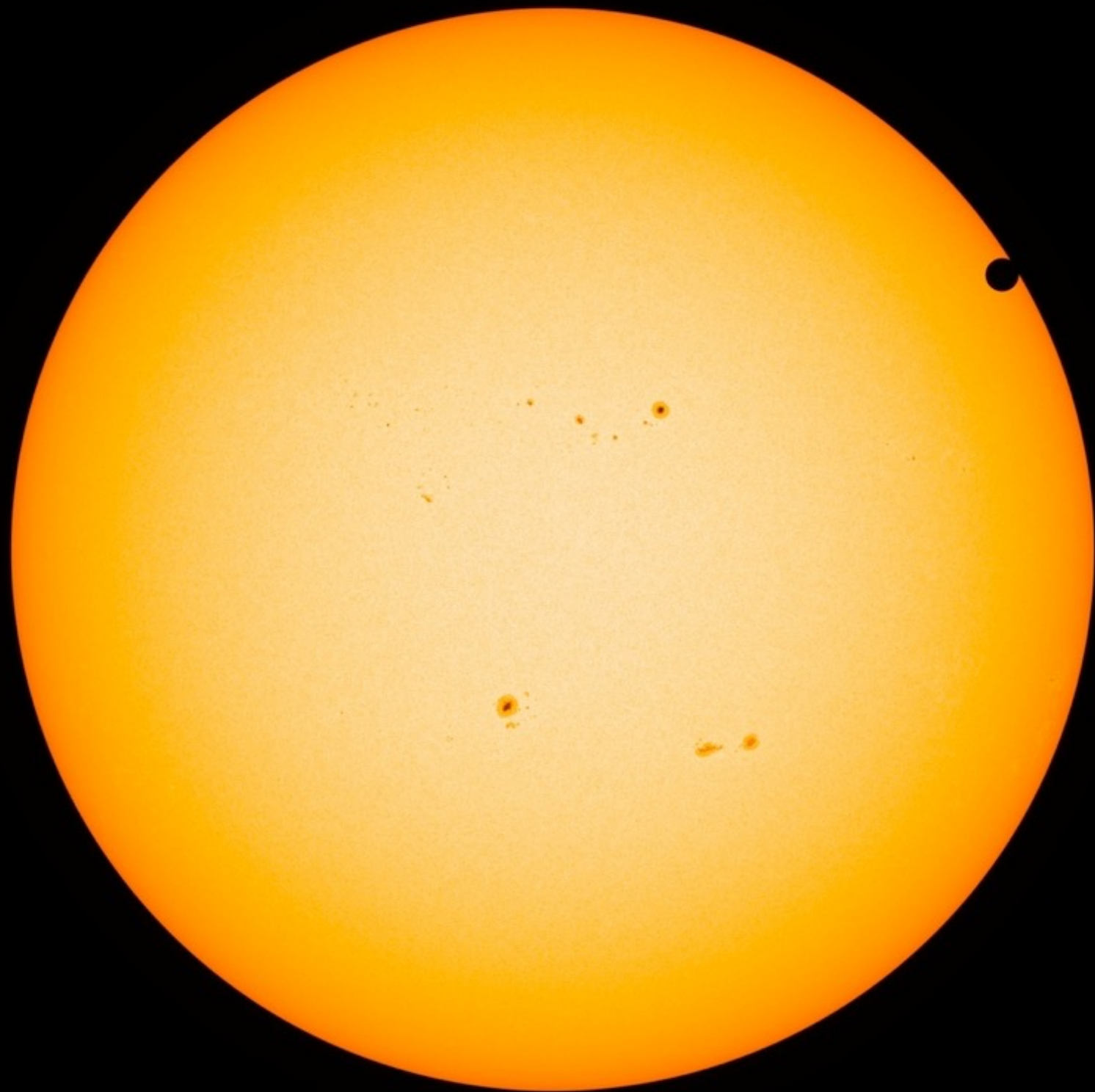
Atmospheric Transmission and Near-IR Filters

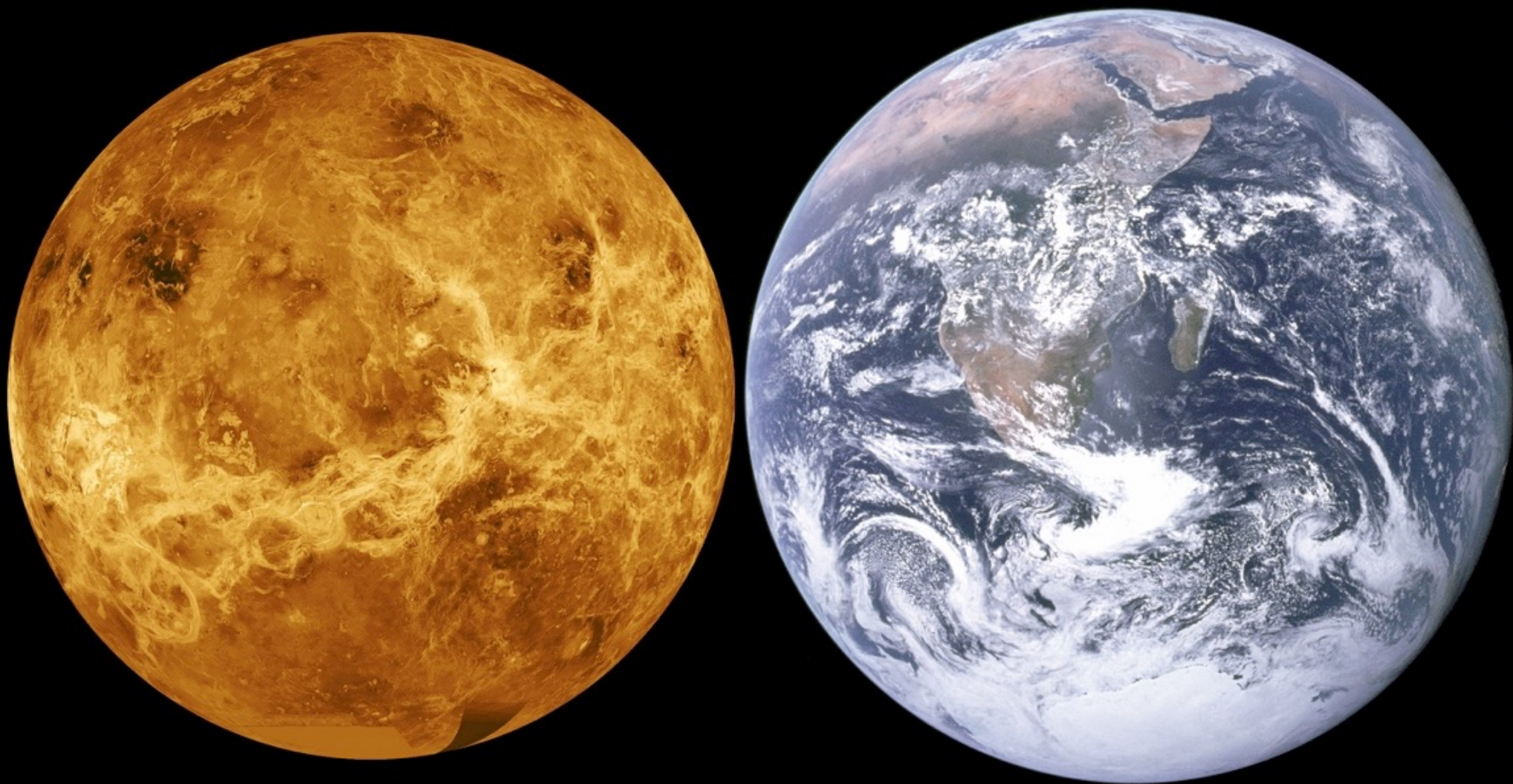


Vega Spectrum











# Sun's Spectrum vs. Thermal Radiator

of a single temperature  $T = 5777 \text{ K}$

