

Figure 1: The three flat, pressureless cosmological models. On the left $\Lambda > 0$, in the middle $\Lambda < 0$ and on the right the Einstein de-Sitter model with $\Lambda = 0$.

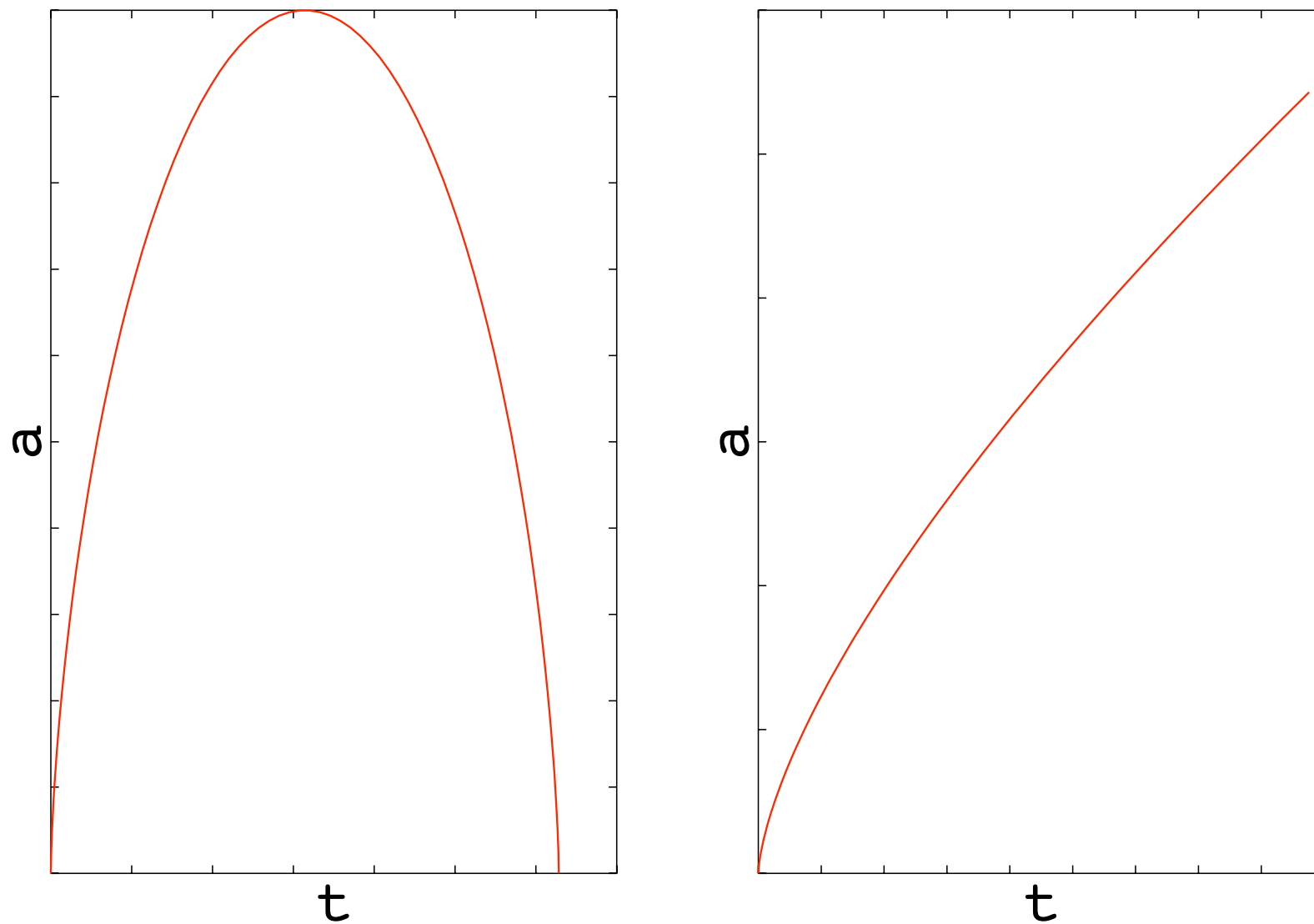


Figure 1: The two non-flat FRW models with vanishing cosmological constant $\Lambda = 0$. On the left with $k = +1$ and on the right with $k = -1$.

