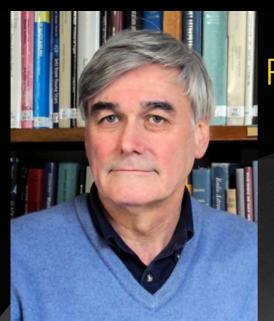
## Dr. Giorgia Busso Institute of Astronomy, Cambridge University

King's Lynn and District Astronomical Society

April 3 2017



## Prof. Gerry Gilmore

#### Dr. Anna Hourihane

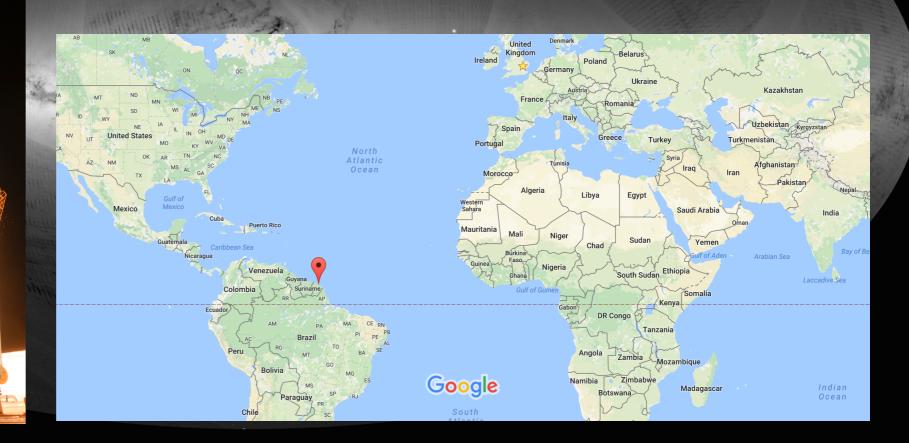


Gaia Team at IoA

Gaia: the billion star surveyor ESA corner mission 1% the visible Milky Way Unprecedented accuracy (µMag and mas)

# **The Mission**

Launched on December 19th, 2013 at 09:12:19 UTC from Kourou in French Guiana.



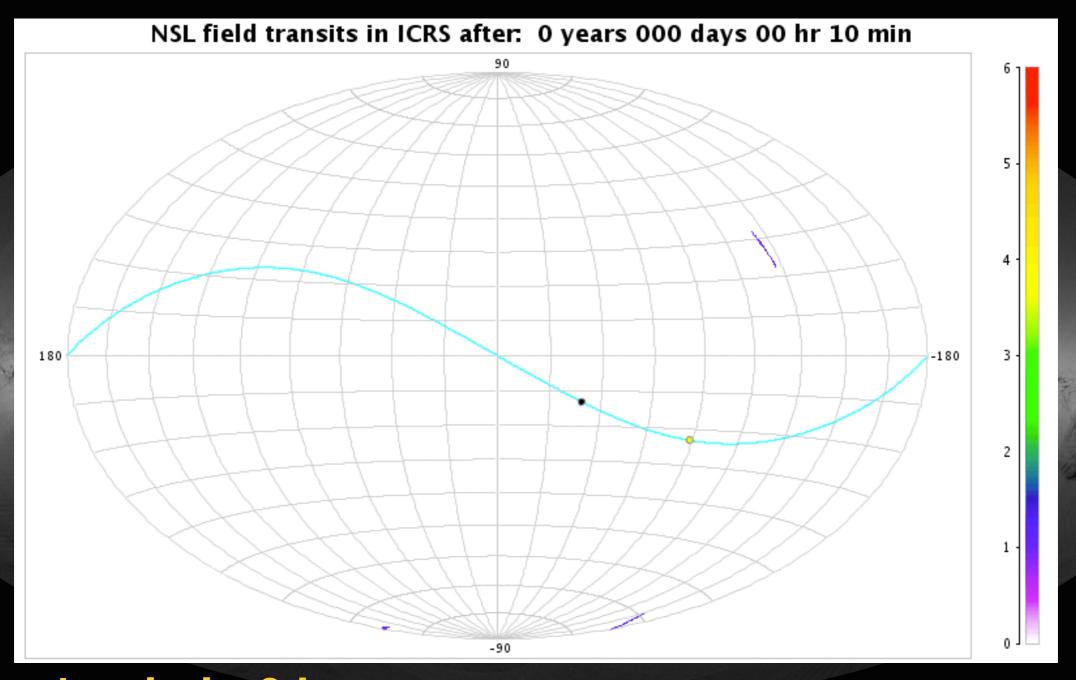
King's Lynn, April 3 2017





King's Lynn, April 3 2017

# **The Mission**



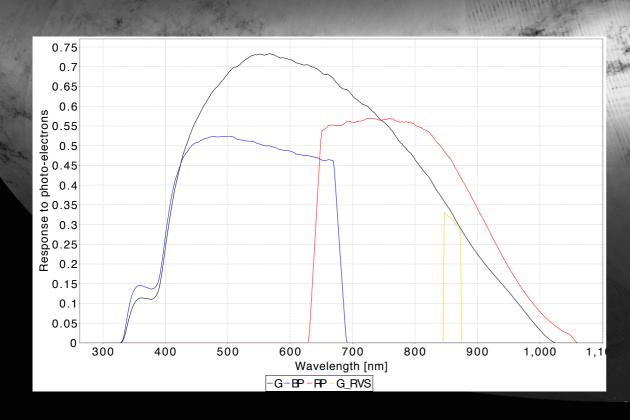
# 1 spin in 6 hours Whole Sky coverage in ~6 months In average a star is observed ~80 times

King's Lynn, April 3 2017

# On board:

# **The Mission**

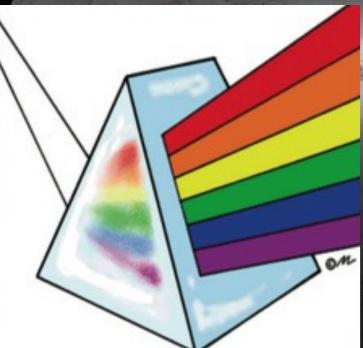
- 2 telescopes with 1.45 x 0.5 m primary mirrors
- Astrometric Field (White Light)
- Blue And Red Photometers
- High-Res Spectrometer

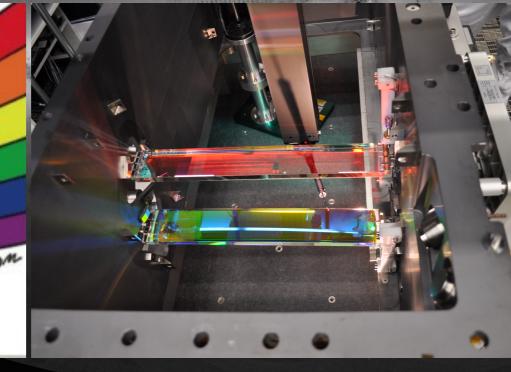


# On board:

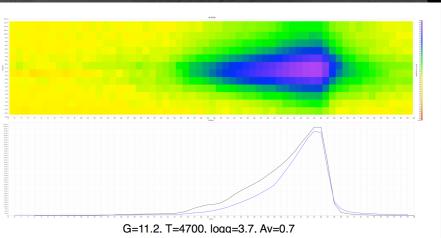
# **The Mission**

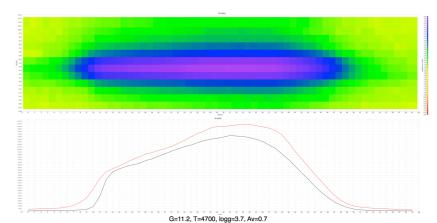
- 2 telescopes with 1.45 x 0.5 m primary mirrors
- Astrometric Field (White Light)
- Blue And Red Photometers
- High-Res Spectrometer





**BP/RP** 





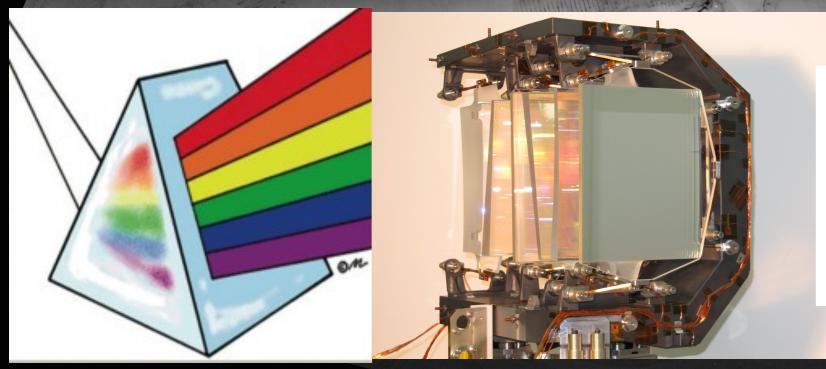
Light dispersion

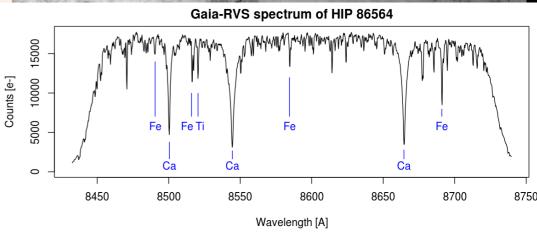


# On board:

# **The Mission**

- 2 telescopes with 1.45 x 0.5 m primary mirrors
- Astrometric Field (White Light)
- Blue And Red Photometers
- High-Res Spectrometer





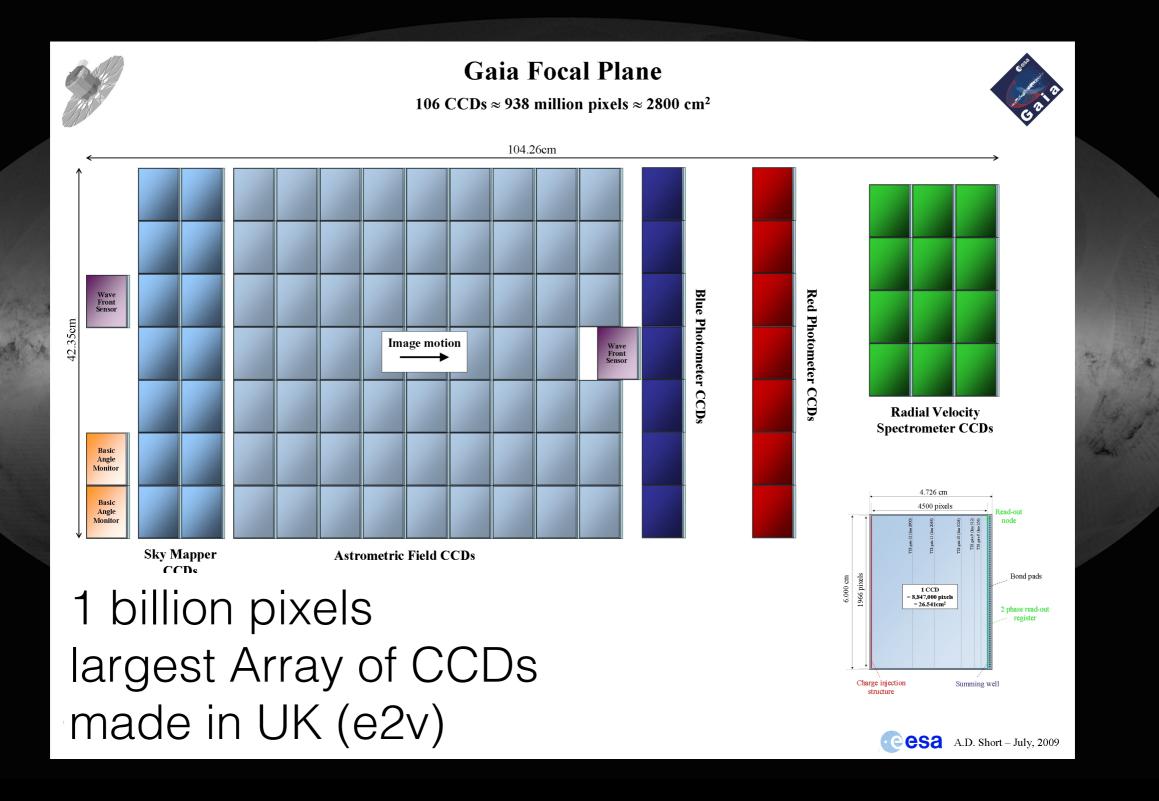
# Light dispersion

RVS



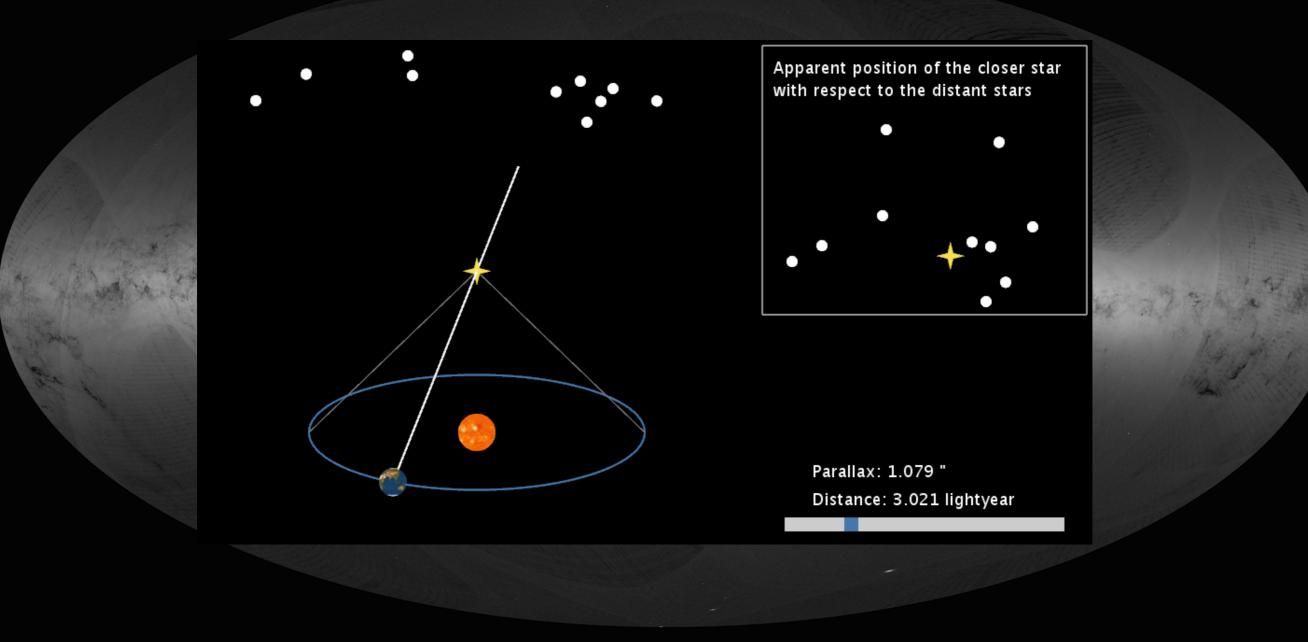
# On board:

# **The Mission**



# **The Mission**

## **Measure Position and Parallax to get distances**



# Not only distances but also..

# **The Mission**



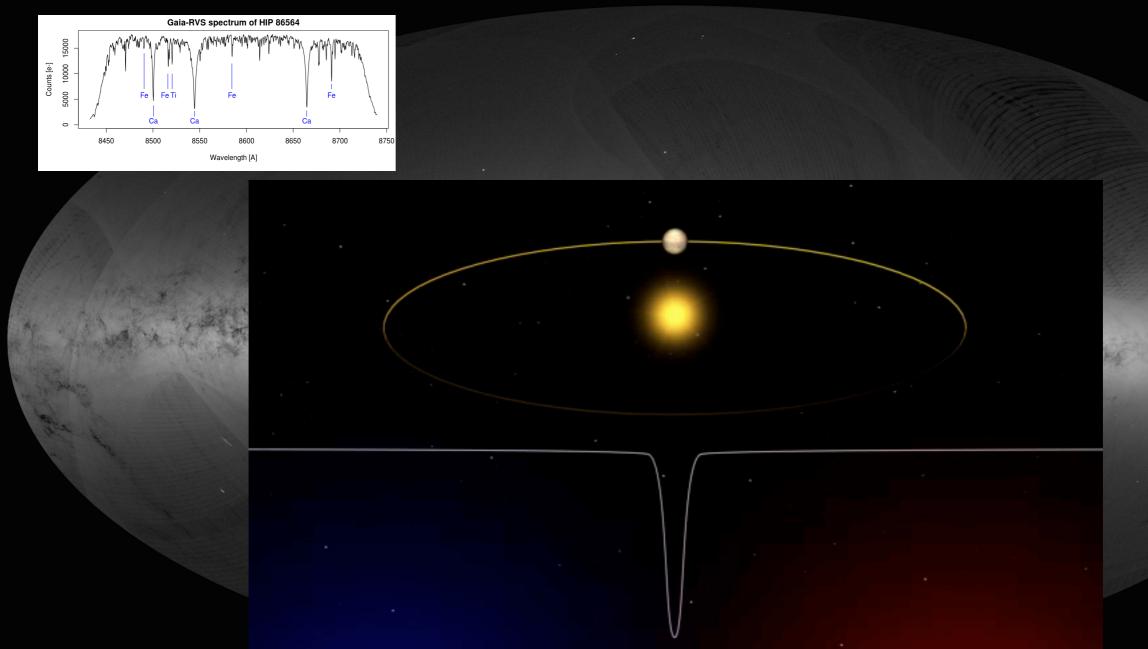
**Proper Motions**,

Apparent motion of a star seen from Earth



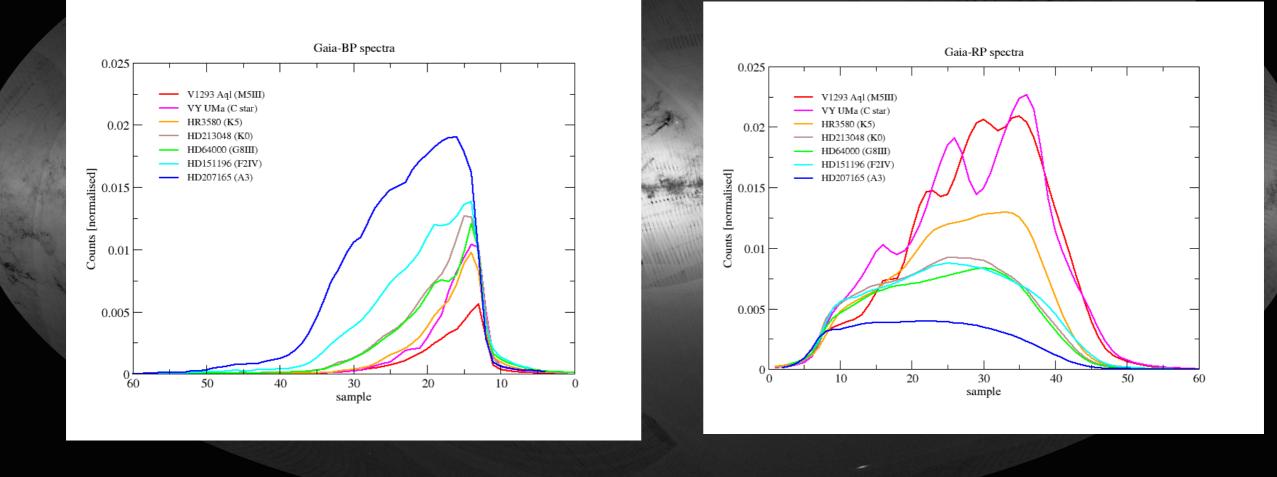
# **Radial Velocities,**

# **The Mission**



#### Credit: ESO/H. Zodet

# and Astrophysical Parameters! Temperature, Chemical Composition, Surface Gravity

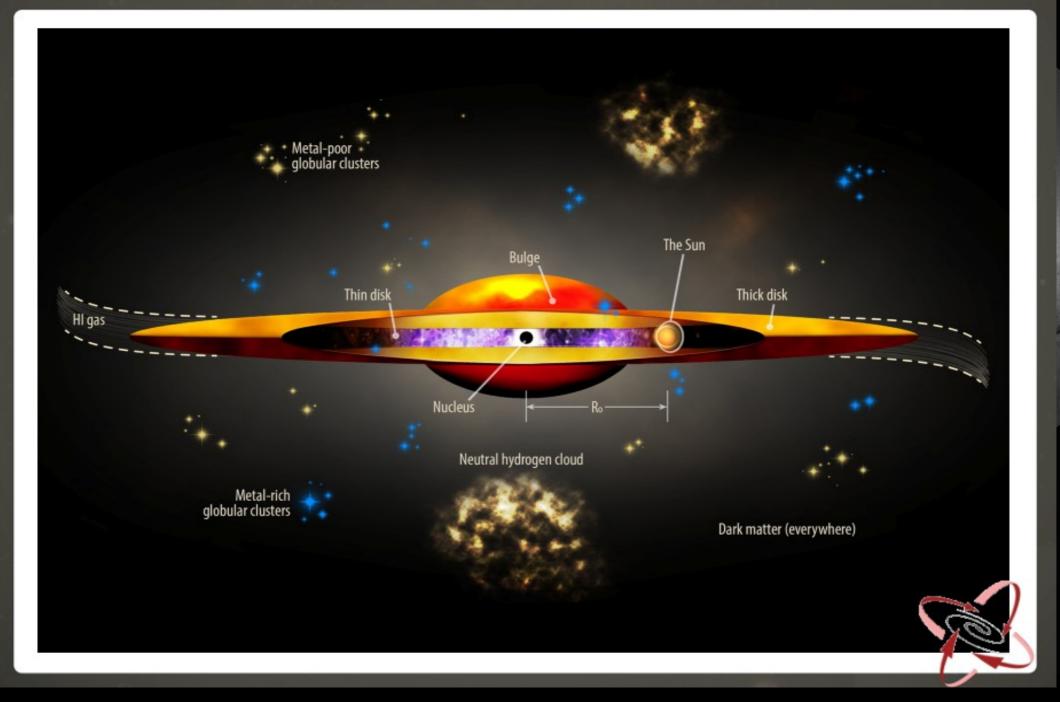


**The Mission** 

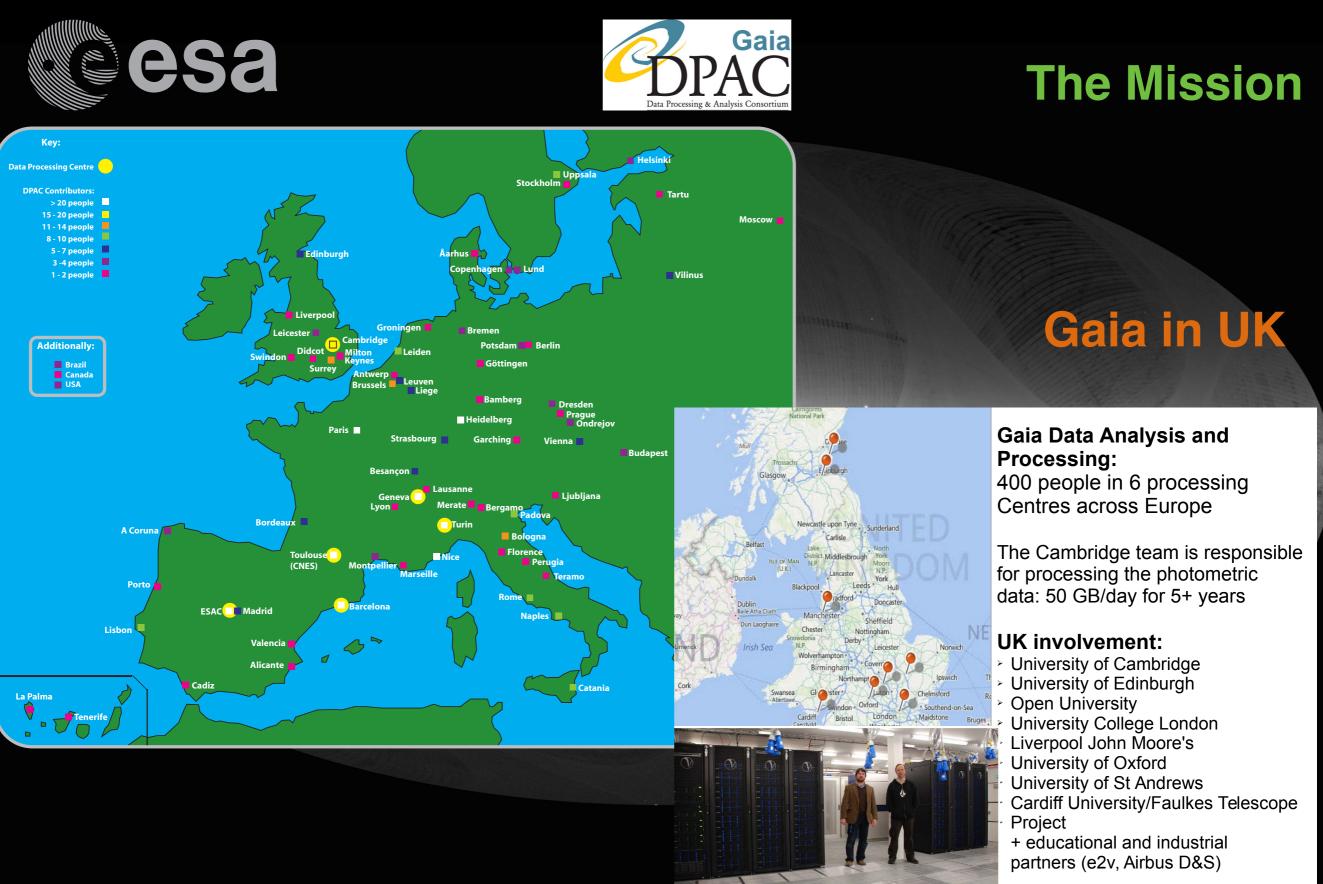
# **The Mission**

#### Putting it all together...

#### Taking the census of the Milky Way Galaxy



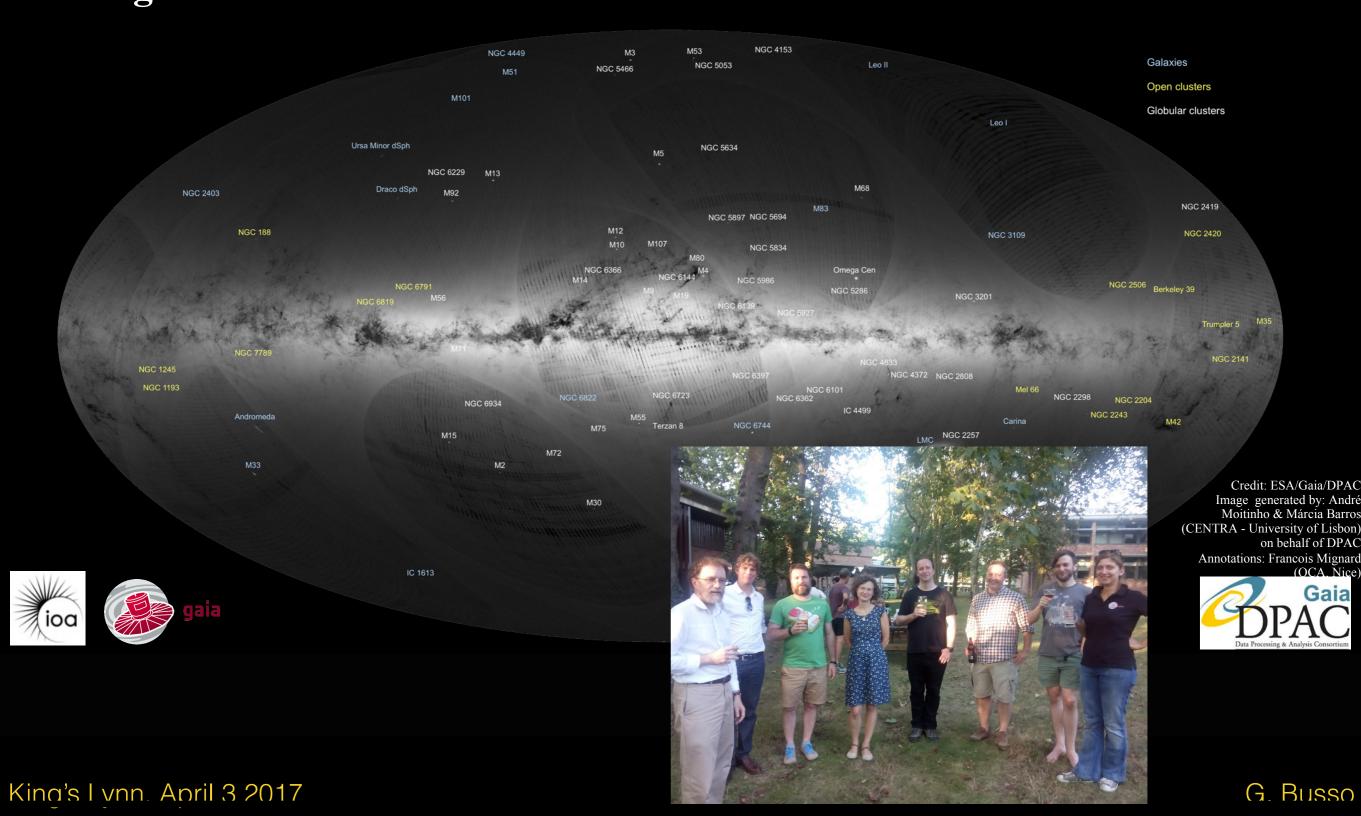
King's Lynn, April 3 2017

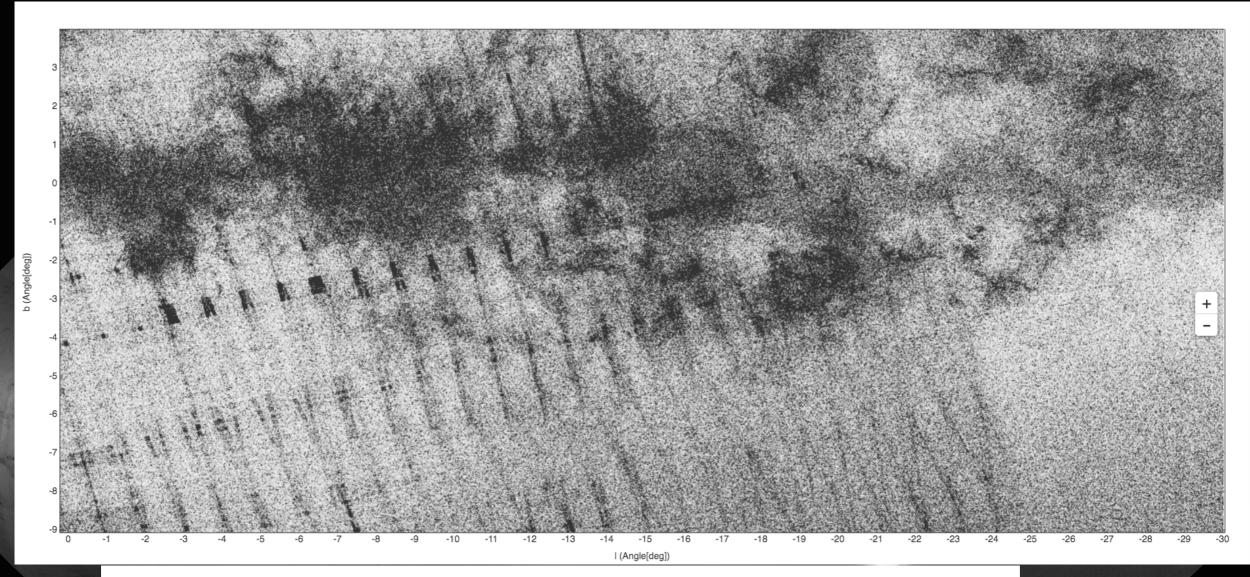


# September 14th, 2016

# Gaia DR1 Sky: all sky high resolution image shows Gaia source densities







Scanning pattern is visible
dense regions
areas with poor scanning law coverage
filtering on number of observations

## → GAIA DATA RELEASE 1

14 September 2016

111111111111

1000 days since launch 2 telescopes 10 mirrors l camera 106 CCDs 937,782,000 pixels

1 spacecraft

500,000 km from

#### Data challenge so far

- >50 billion focal plane transits
- >110 billion photometric observations
- >9.4 billion spectroscopic observations
- ~120,000 hours of computing time to identify stars
- 6 data processing centres

## **1 Milky Way**

>100,000,000,000 stars ~13,000,000,000 years old

#### **Content of the release**

Total number of sources in primary astrometric data set: 2,057,050

with position, magnitude, parallax & proper motion

Total number of sources in secondary astrometric data set: 1,140,622,719 with position & magnitude

3194 Variable stars

 599 Cepheids (43 new discoveries) 2595 RR Lyrae (343 new discoveries)

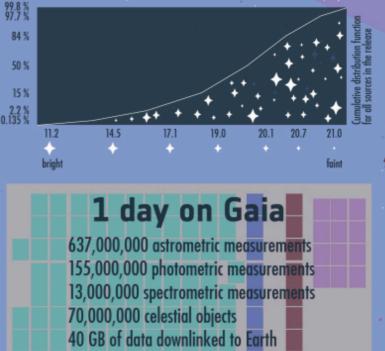
#### 2152 Quasars

with position & magnitude

Data collected over 14 months

# Magnitude distribution

esa



Intermediate release -> Only positions and some proper motions **No Colors or Radial Velocities** 

King's Lynn, April 3 2017

www.esa.int

European Space Agency

# Precisions: in mas (but will be in µas at the end)

#### ASTROMETRIC UNCERTAINTY

#### **TGAS - astrometric uncertainties**

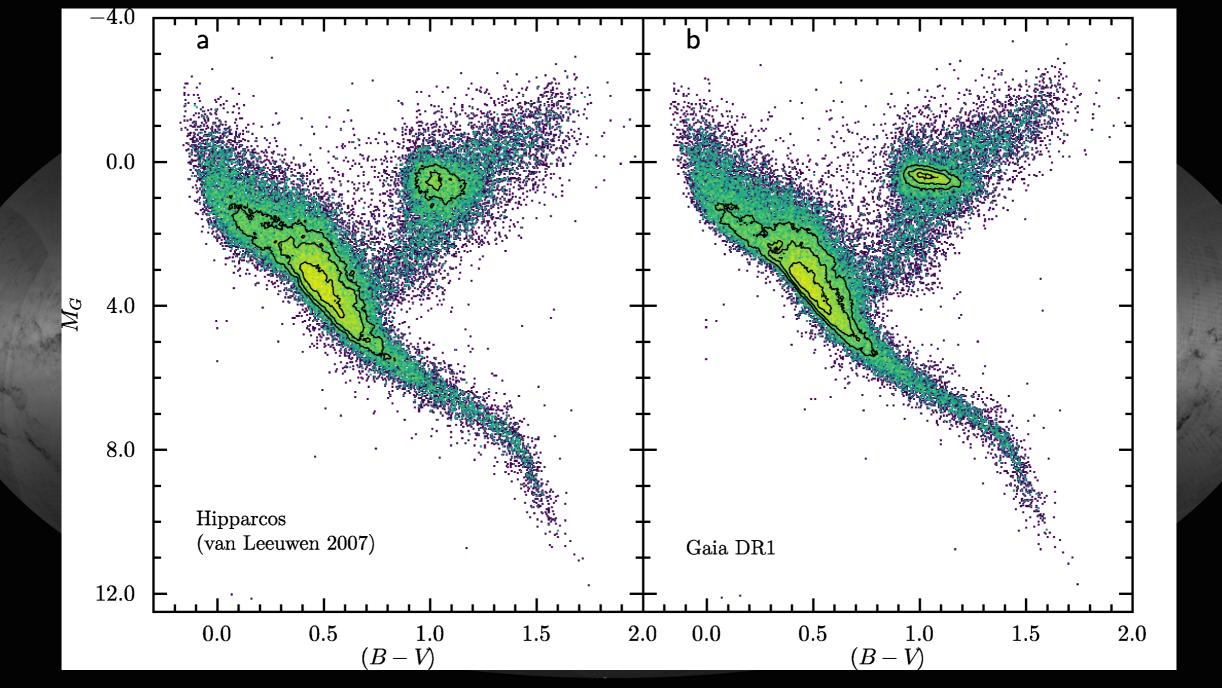
	All TGAS sources			Hipparcos subset		
	10%	50%	90%	10%	50%	90%
G magnitude	9.29	11.04	12.05	7.00	8.32	9.73
position (mas)	0.20	0.32	0.75	0.20	0.26	0.46
parallax (mas)	0.24	0.32	0.64	0.23	0.28	0.48
proper motion (mas yr <sup>-1</sup> )	0.72	1.32	3.19	0.04	0.07	0.14

These are precisions. There is an additional systematic parallax error of  $\pm 0.3$  mas.

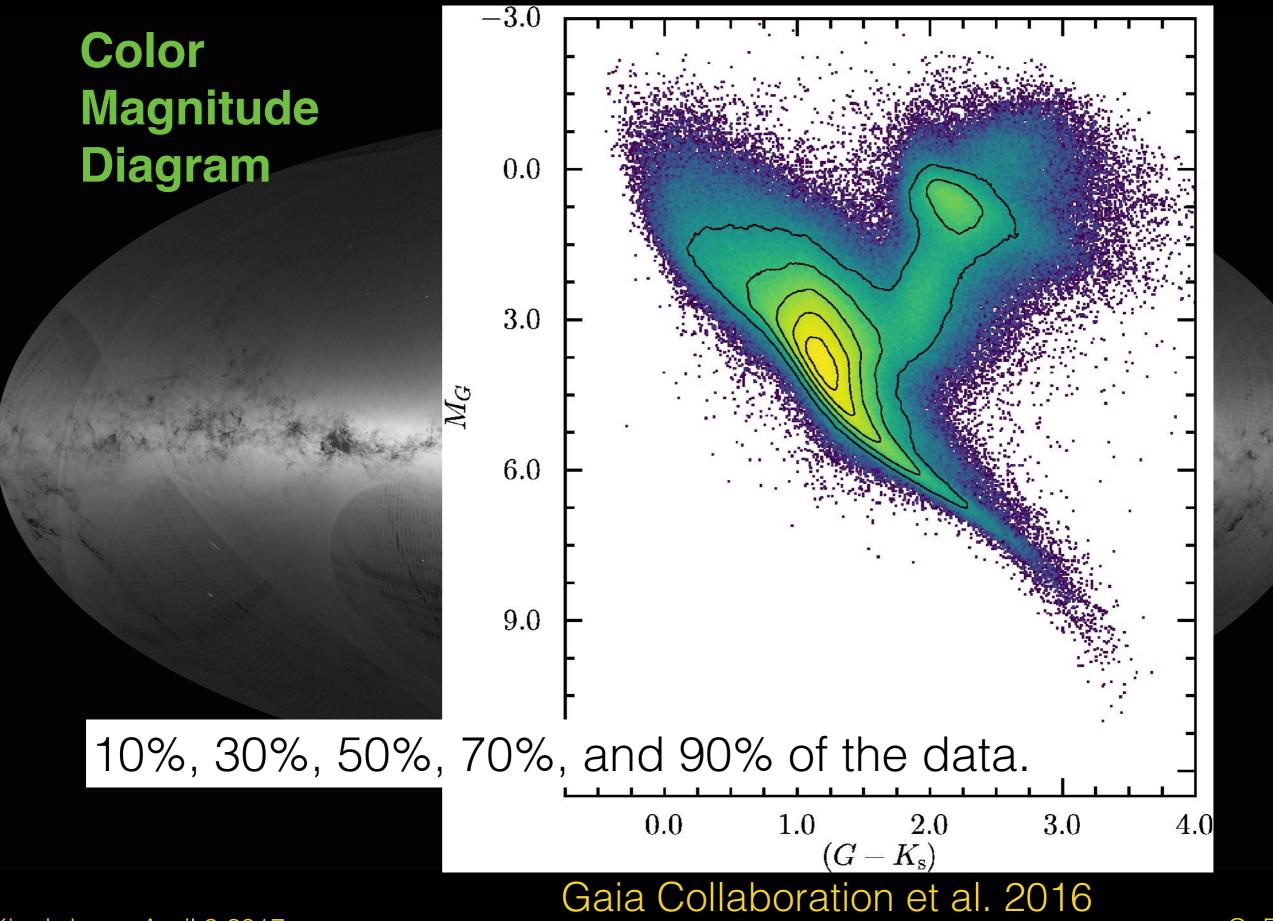
#### Secondary set - positional uncertainties

G mag	fraction	10%	50%	90%
< 16	7%	0.1	0.3	5.3
16-17	7%	0.2	0.5	12.1
17-18	12%	0.3	0.8	12.4
18-19	21%	0.5	1.5	13.7
19-20	30%	0.9	2.7	16.6
20-21	22%	1.9	2.4	21.5
All	100%	0.35	2.4	16.3
Quantile positional precisions	in mas.		•	•

# **Comparison with Hipparcos**



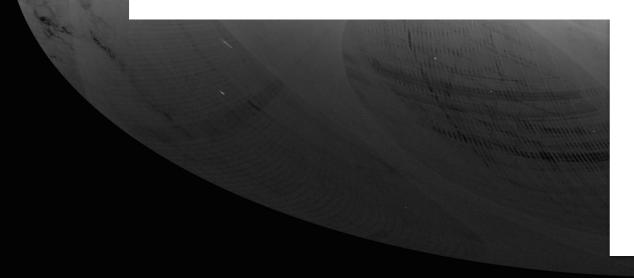
Gaia Collaboration et al. 2016

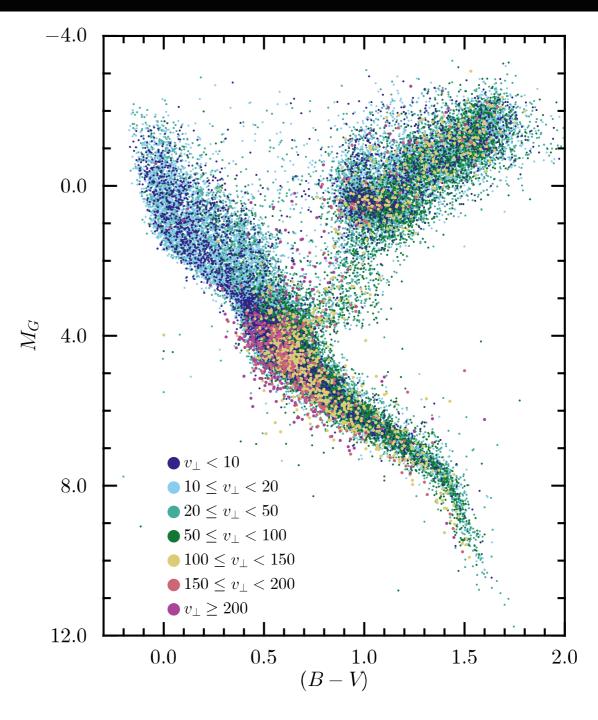


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# Color Magnitude Diagram

# Young disk Old disk (metal rich) Halo (metal poor)

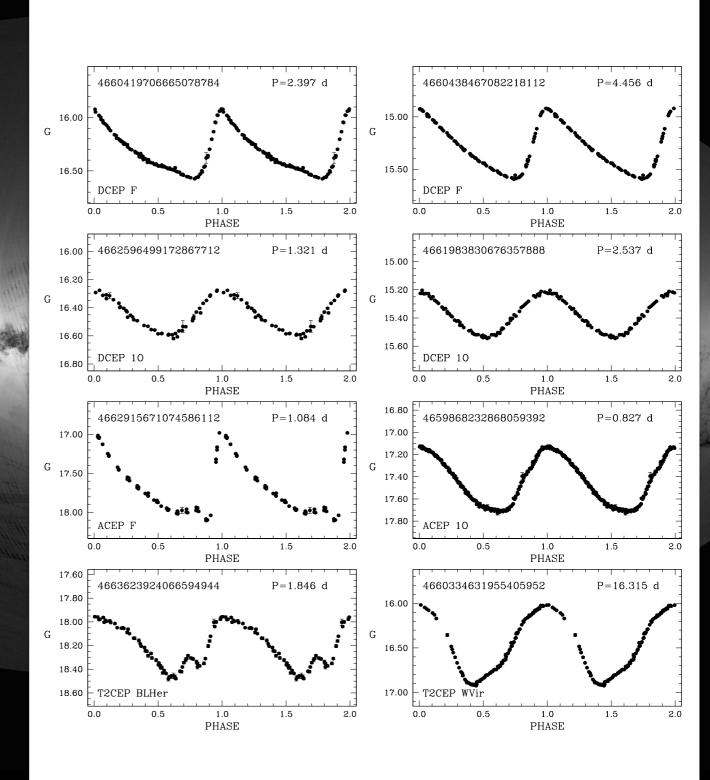




Gaia Collaboration et al. 2016

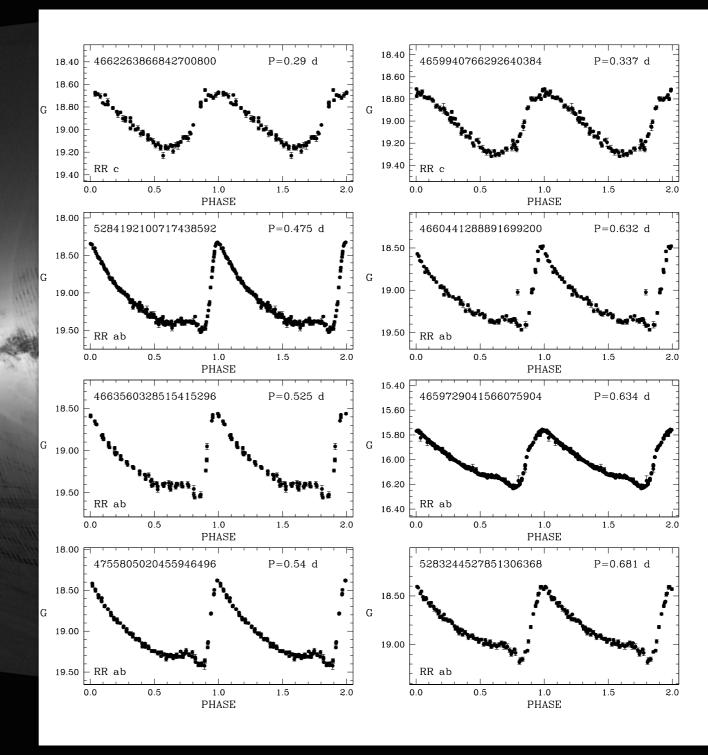
# Variable stars Light Curves

# 599 Cepheid (43 new)



# Variable stars Light Curves

# 2595 RR-Lyrae (343 new)



## The very first paper after GDR1

APJ LETTERS, SUBMITTED, SEPTEMBER 14, 2016 Preprint typeset using  ${}^{\text{ATE}}_{\text{EX}}$  style emulateapj v. 5/2/11

#### FIRST GAIA LOCAL GROUP DYNAMICS: MAGELLANIC CLOUDS PROPER MOTION AND ROTATION

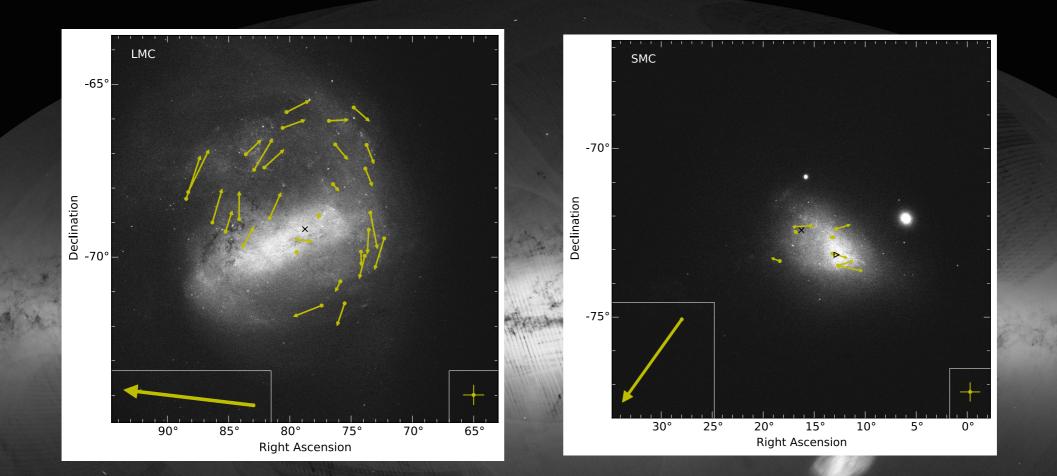
ROELAND P. VAN DER MAREL Space Telescope Science Institute, 3700 San Martin Drive, Baltimore, MD 21218

JOHANNES SAHLMANN European Space Agency, Space Telescope Science Institute, 3700 San Martin Drive, Baltimore, MD 21218, USA ApJ Letters, submitted, September 14, 2016

#### ABSTRACT

We use the *Gaia* data release 1 (DR1) to study the proper motion (PM) fields of the Large and Small Magellanic Clouds (LMC, SMC). This uses the *Tycho-Gaia* Astrometric Solution (TGAS) PMs for 29 *Hipparcos* stars in the LMC and 8 in the SMC. The LMC PM in the West and North directions is inferred to be  $(\mu_W, \mu_N) = (-1.872 \pm 0.045, 0.224 \pm 0.054)$  mas yr<sup>-1</sup>, and the SMC PM  $(\mu_W, \mu_N) = (-0.874 \pm 0.066, -1.229 \pm 0.047)$  mas yr<sup>-1</sup>. These results have similar accuracy and agree to within the uncertainties with existing Hubble Space Telescope (*HST*) PM measurements.

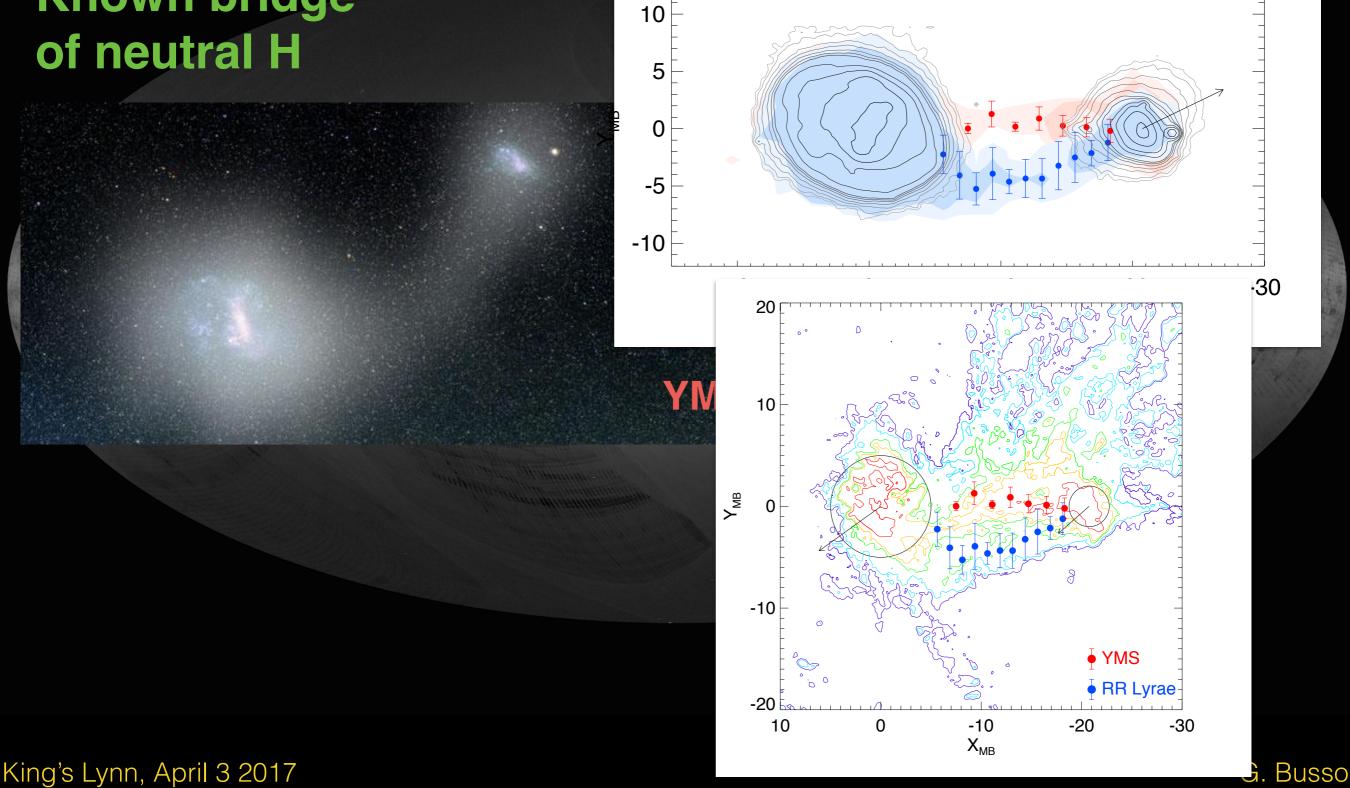
# The very first paper after GDR1



# Confirm of rotation of Large and Small Magellanic Clouds

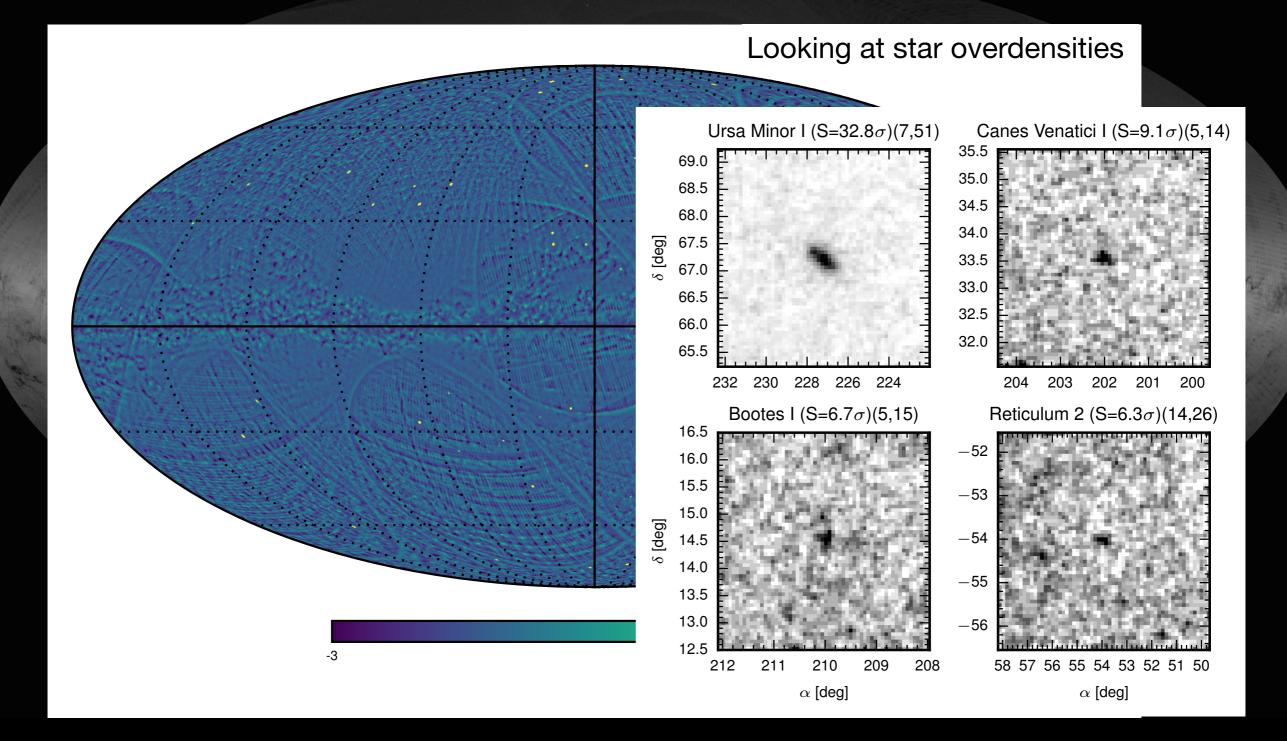
# **Magellanic Clouds, Streams and Bridges**

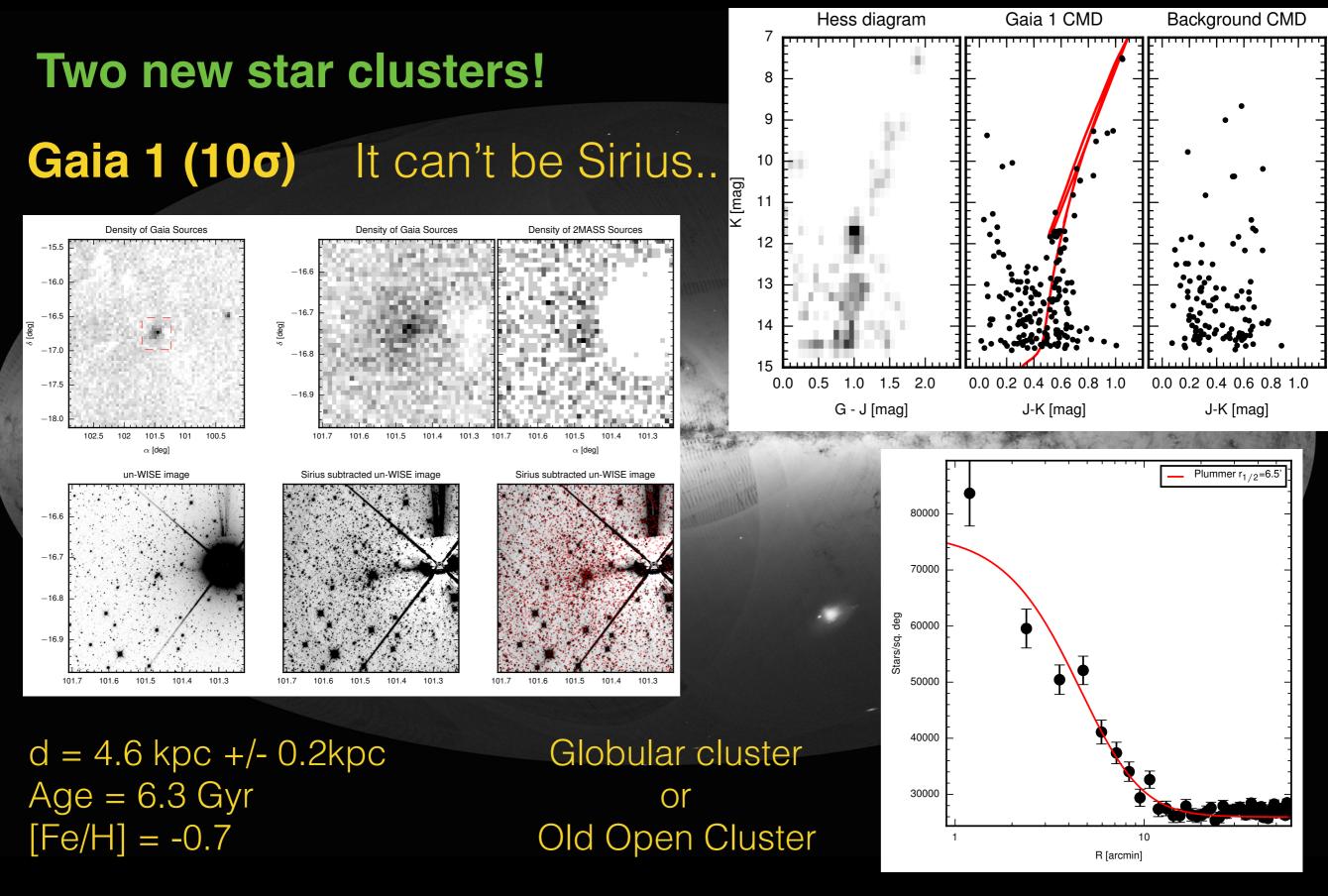
# **Known bridge**



## **Two new star clusters!**

(Koposov, Belokurov and Torrealba, IoA, Cambrigde)

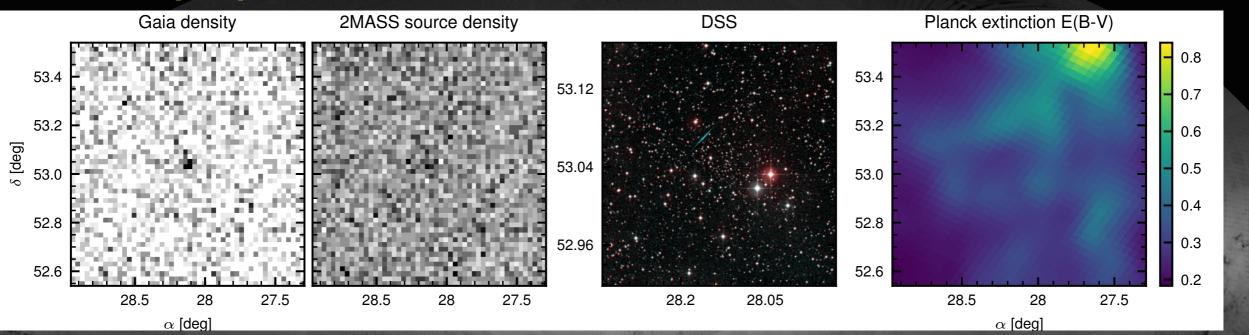


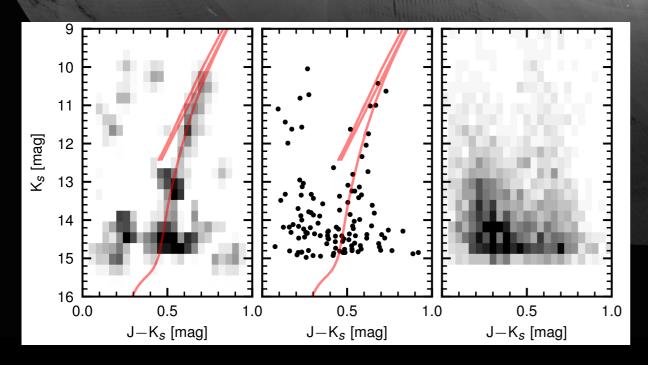


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# **Two new star clusters!**

# Gaia 2 (9σ)





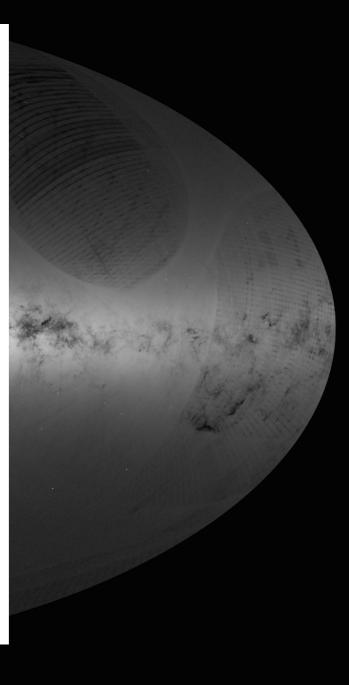
d = 5.5 kpc +/- 0.2kpc Age = 10 Gyr [Fe/H] = -1

Globular cluster or Old Open Cluster

## **Science Alerts**

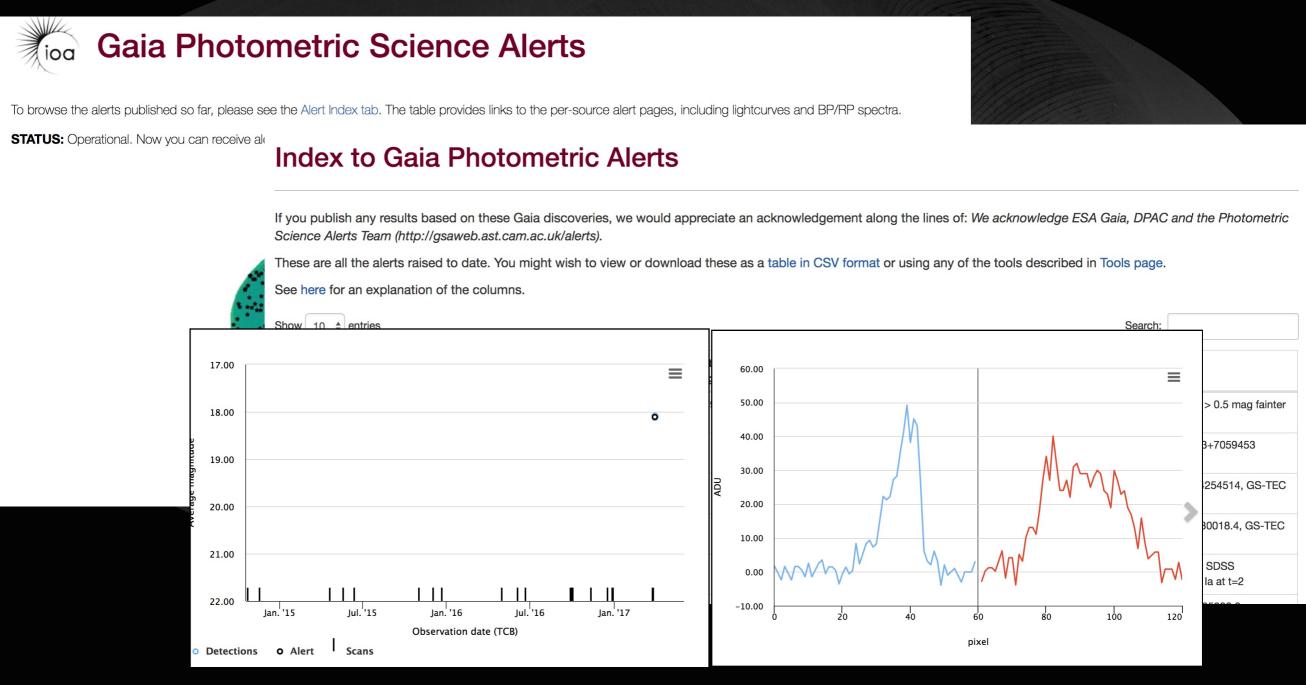
# Promptness of publication

- Upstream processing delivers data ~24+ hours after observation, roughly one run per day
- Alerts processing (light-curve assembly, calibration, transient detection and classification) takes up to 6 hours per run
- Publication latency after alerts processing:
  - If classification & selection is automatic: ~ minutes
  - If classification & selection is manual: ~ hours to ~ days

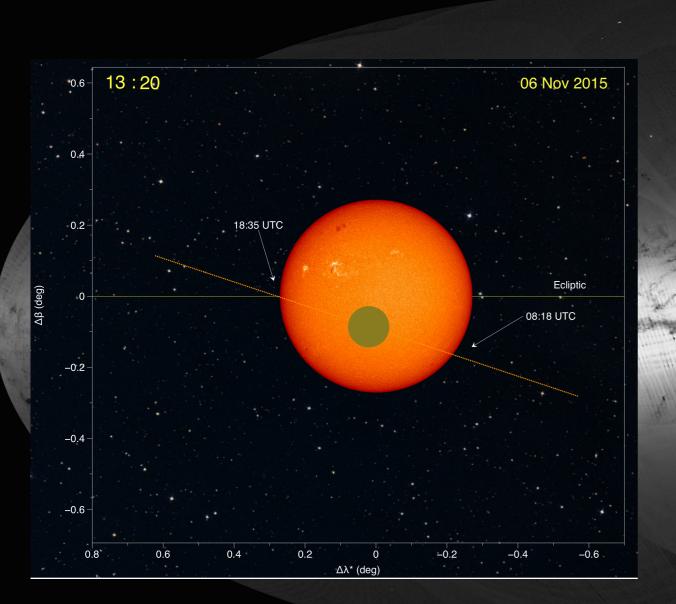


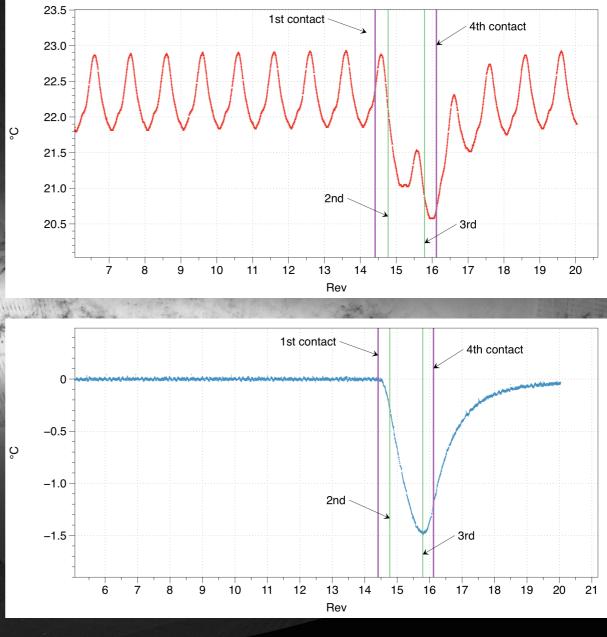
# **Science Alerts**

# http://gsaweb.ast.cam.ac.uk/alerts/home



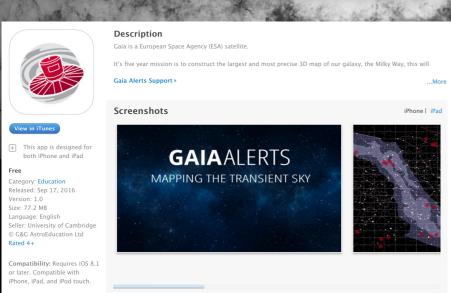
# **Another way of looking at Moon Transits**





# If you want to follow Gaia on your phone:

# https://www.cosmos.esa.int/web/gaia/gaia-app





# https://itunes.apple.com/us/app/gaia-alerts/id1144470584

Android soon on GooglePlay



# **GDR2**

- 5-parameter astrometric solutions for all sources (positions, parallax, proper motions);
- G and integrated GBP and GRP photometric fluxes and magnitudes for all sources. -> colors!
- Median radial velocities for sources brighter than GRVS=12 mag.
- For stars G<17 mag T\_eff and line-of-sight extinction
- Photometric data for a sample of variable stars.
- Epoch astrometry for a pre-selected list of >10,000 asteroids

# **!!April 2018!!**

# stay tuned ...

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				CUPTRIGHT 2017 @ EURUPEAN SPACE A	GENCY. ALL RIGHTS RESERVED. (v1.2.2)	

# https://gea.esac.esa.int/archive/

# GaiaSky



# https://zah.uni-heidelberg.de/gaia/outreach/gaiasky/